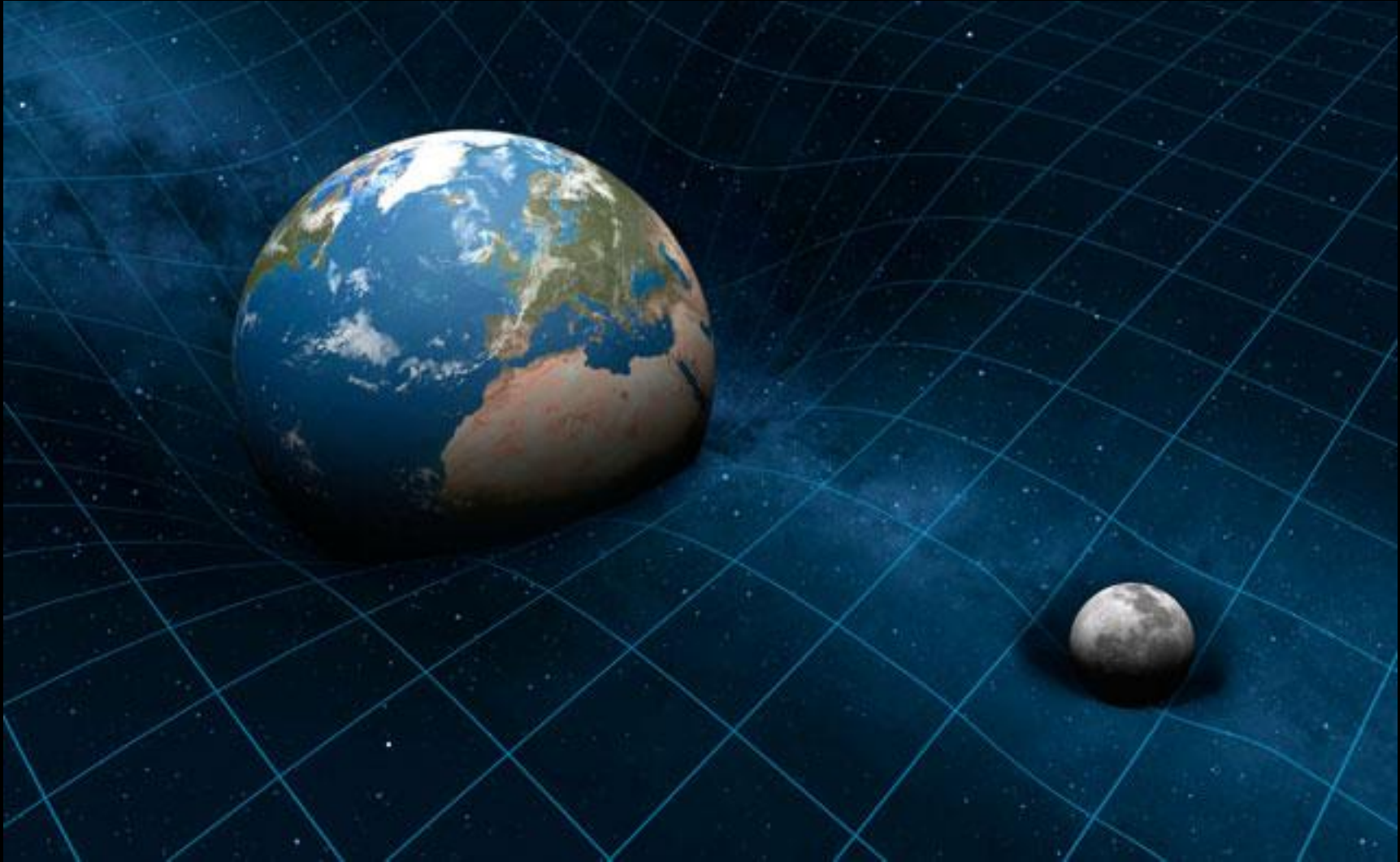
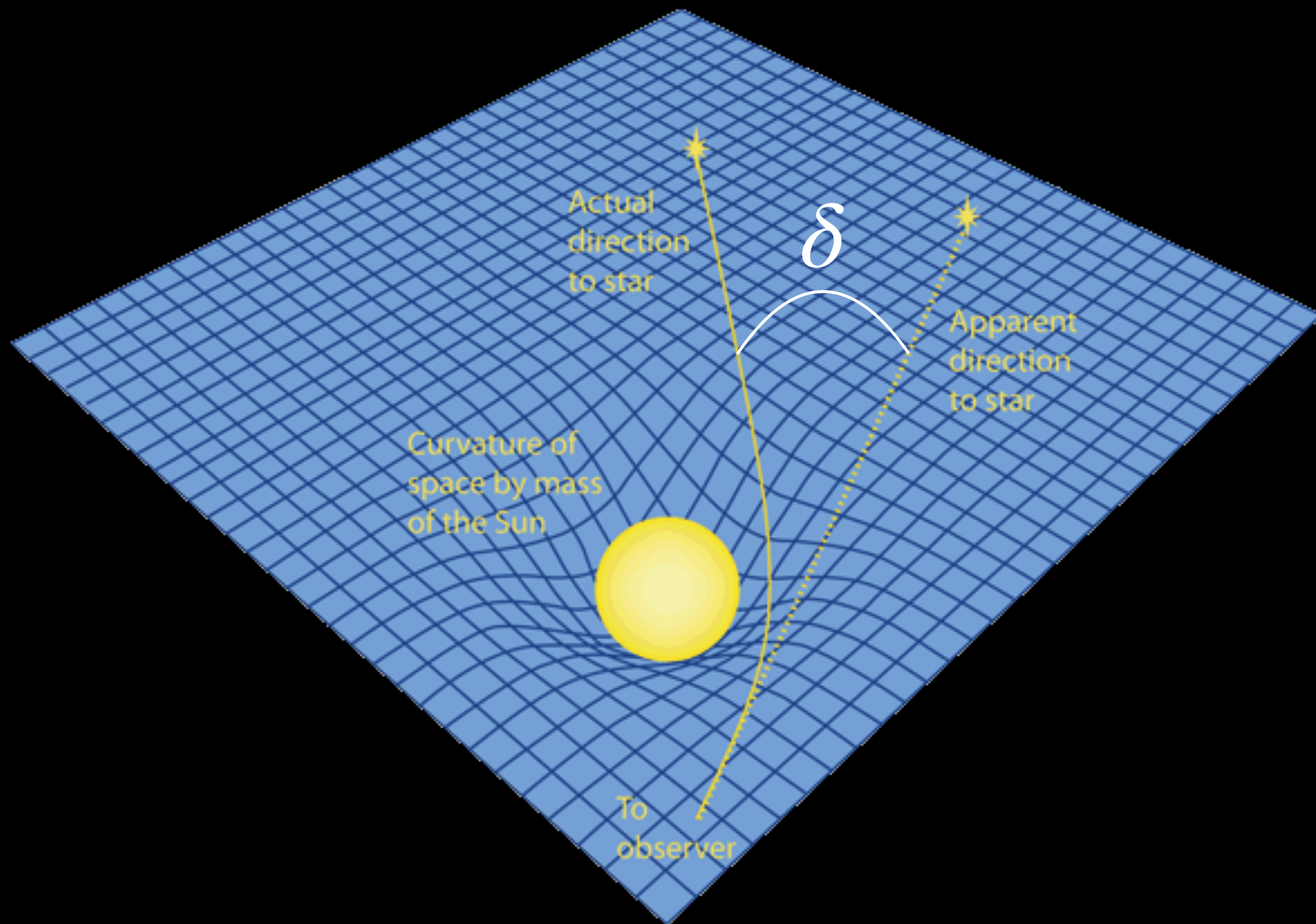


# First Sight of a Black Hole



Recent Results from the  
Event Horizon Telescope





# A Problem

The deflection of starlight is small:

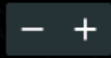
“Newtonian” prediction:  $\delta_N = \frac{2GM_{\text{Sun}}}{c^2 R_{\text{Sun}}} = 0.825 \text{ arcsec}$

Einstein’s prediction:  $\delta_E = 2\delta_N = \frac{4GM_{\text{Sun}}}{c^2 R_{\text{Sun}}}$   
 $= 1.75 \text{ arcsec}$

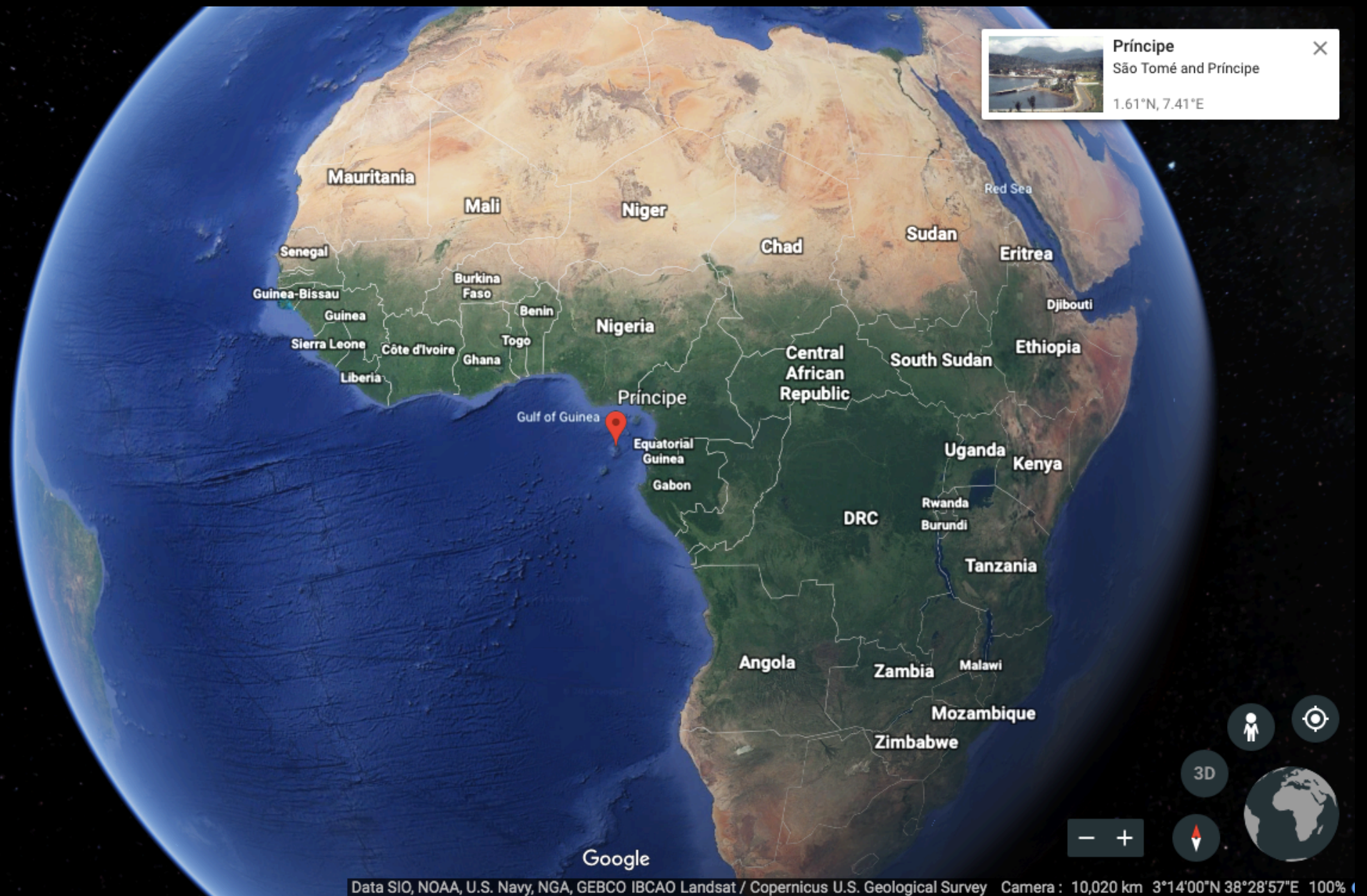





**Sobral**  
State of Ceará  
Brazil  
3.69°S, 40.35°W







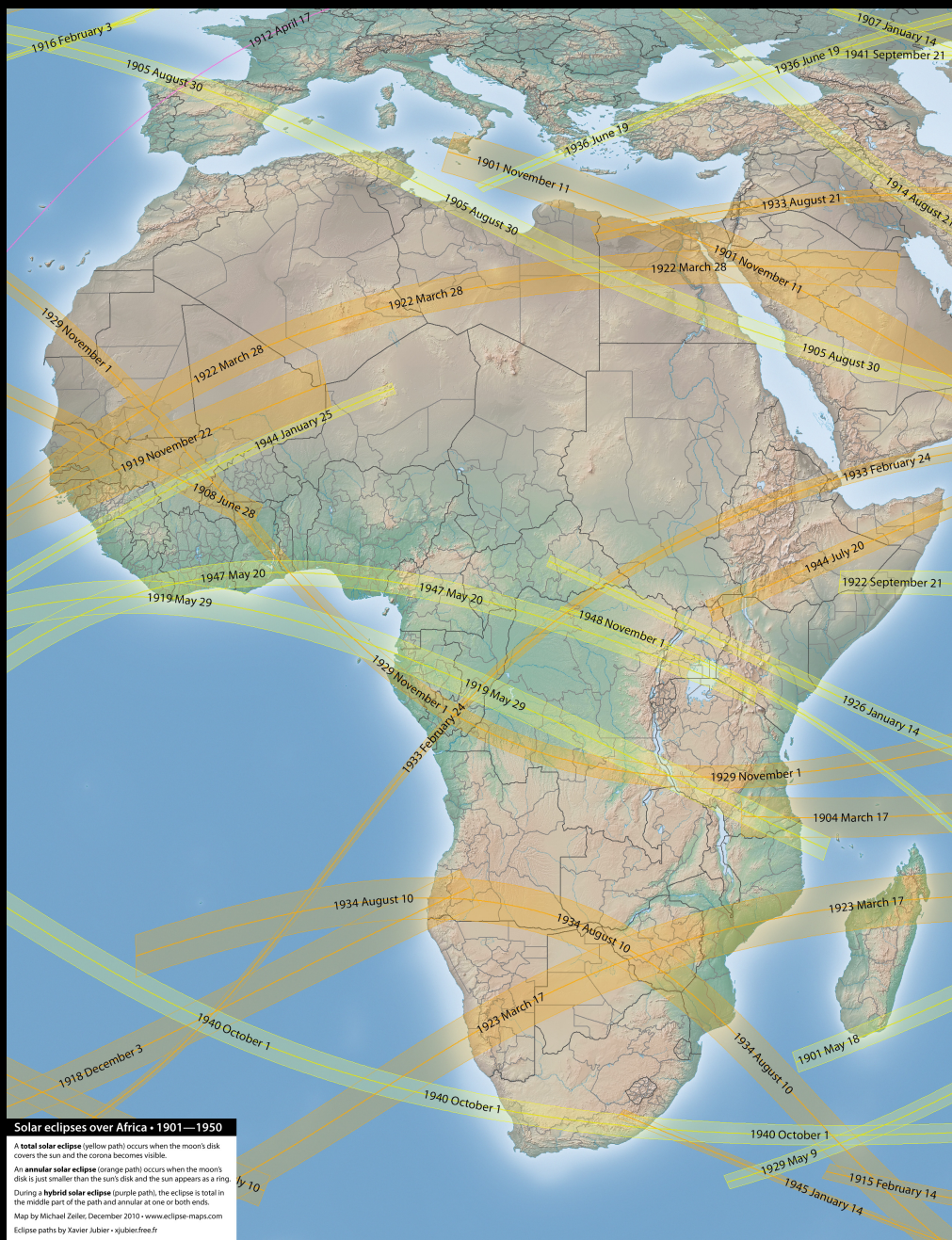
 **Príncipe**  
São Tomé and Príncipe  
1.61°N, 7.41°E

Mauritania  
Mali  
Niger  
Senegal  
Guinea-Bissau  
Guinea  
Sierra Leone  
Côte d'Ivoire  
Liberia  
Burkina Faso  
Benin  
Togo  
Ghana  
Nigeria  
Chad  
Sudan  
Eritrea  
Djibouti  
Ethiopia  
South Sudan  
Central African Republic  
Uganda  
Kenya  
Rwanda  
Burundi  
Tanzania  
Angola  
Zambia  
Malawi  
Mozambique  
Zimbabwe

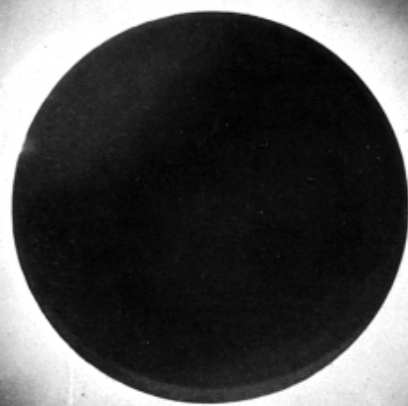
Príncipe  
Equatorial Guinea  
Gabon











# Results

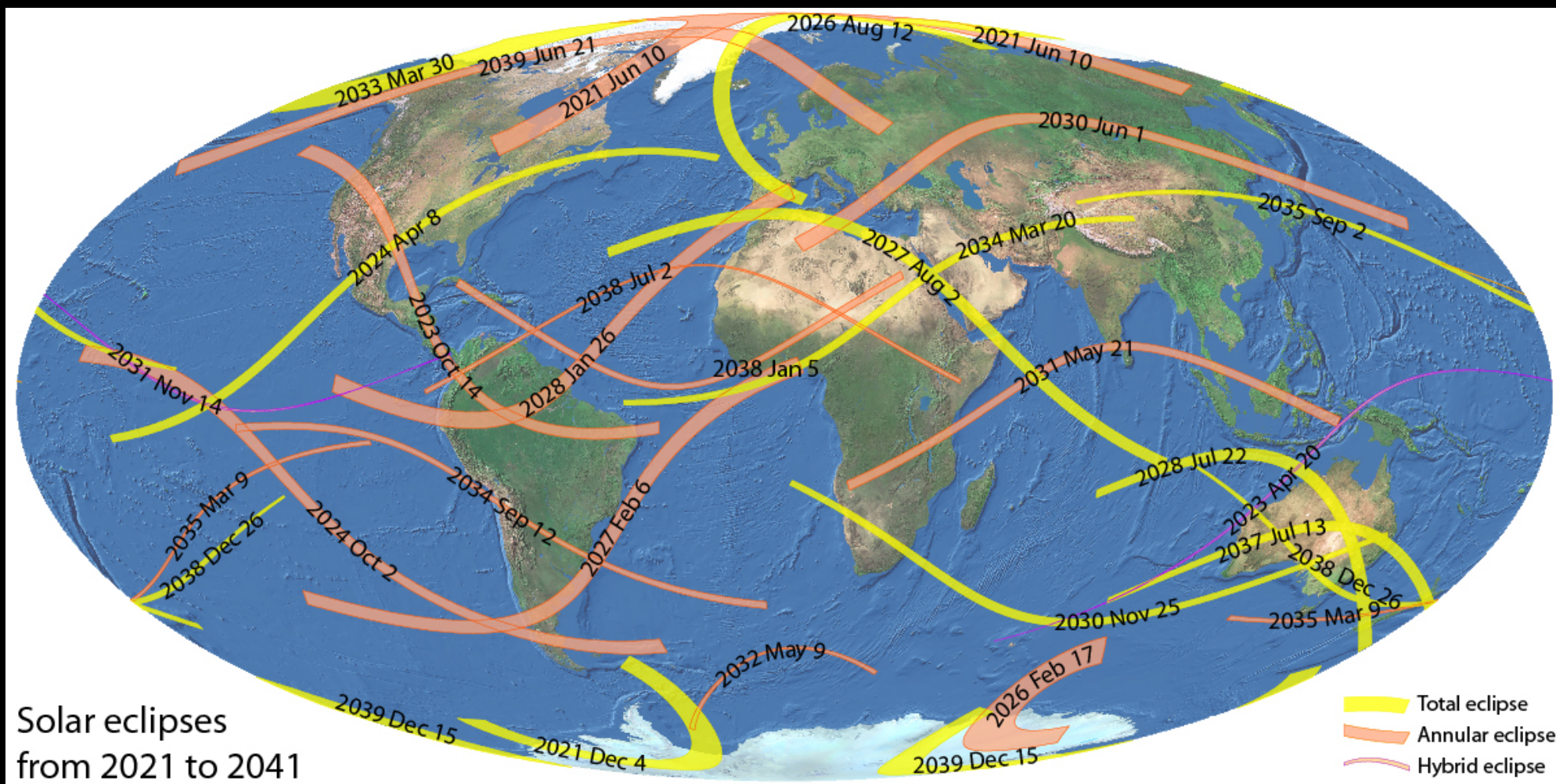
Gravitational displacement at the Sun's limb in seconds of arc

Determination	Displacement
Predicted from Einstein's Theory	1.75
4-inch plates reduced by Dyson <i>et al.</i>	$1.98 \pm 0.18$
4-inch plates measured on the Zeiss	$1.90 \pm 0.11$
Astrographic plates reduced by Dyson <i>et al.</i>	0.93
Astrographic plates measured on the Zeiss	$1.55 \pm 0.34$

Principle result reduced by Eddington  $1.61 \pm 0.3$







# Outline

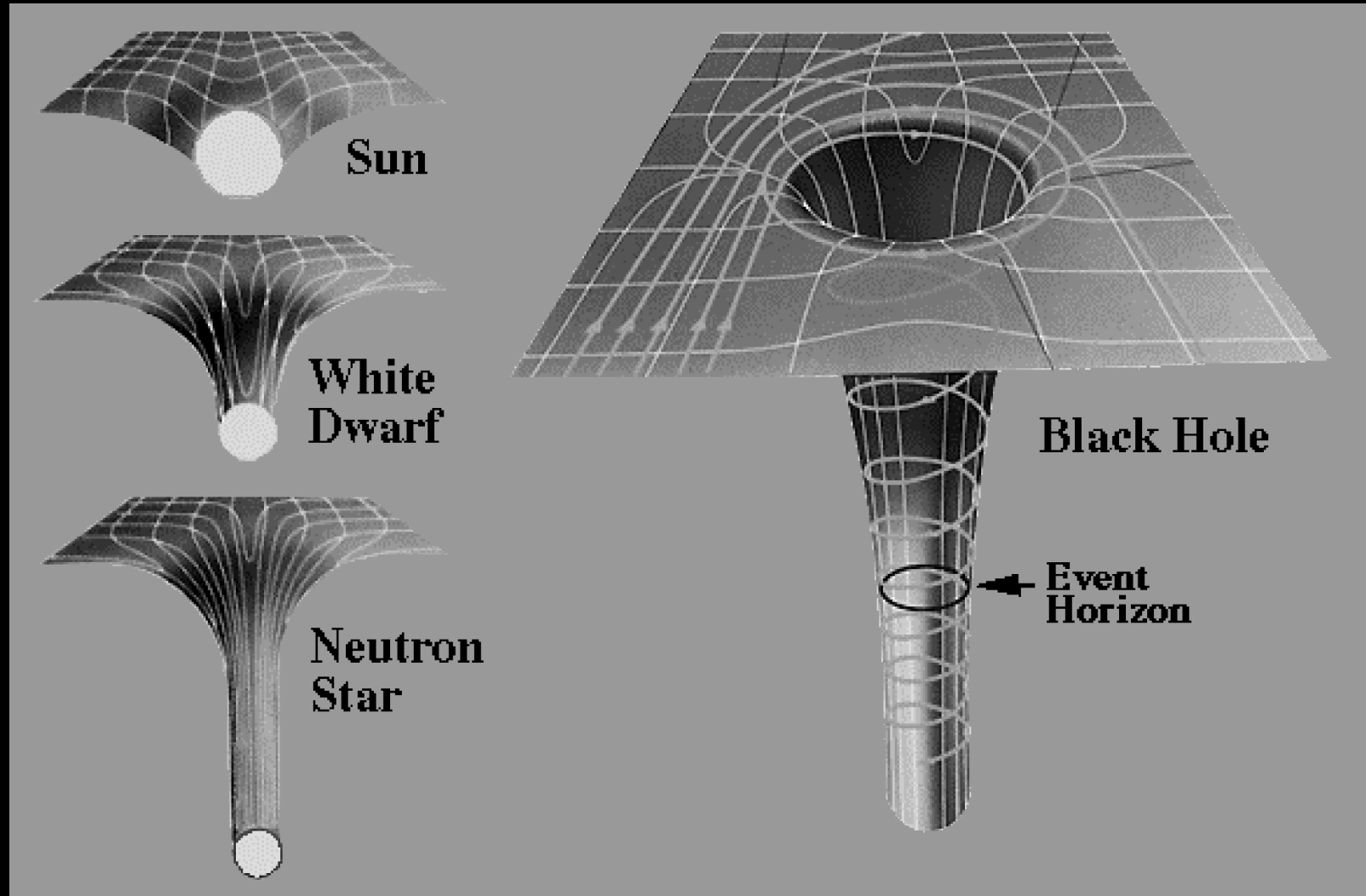
I. Black Holes

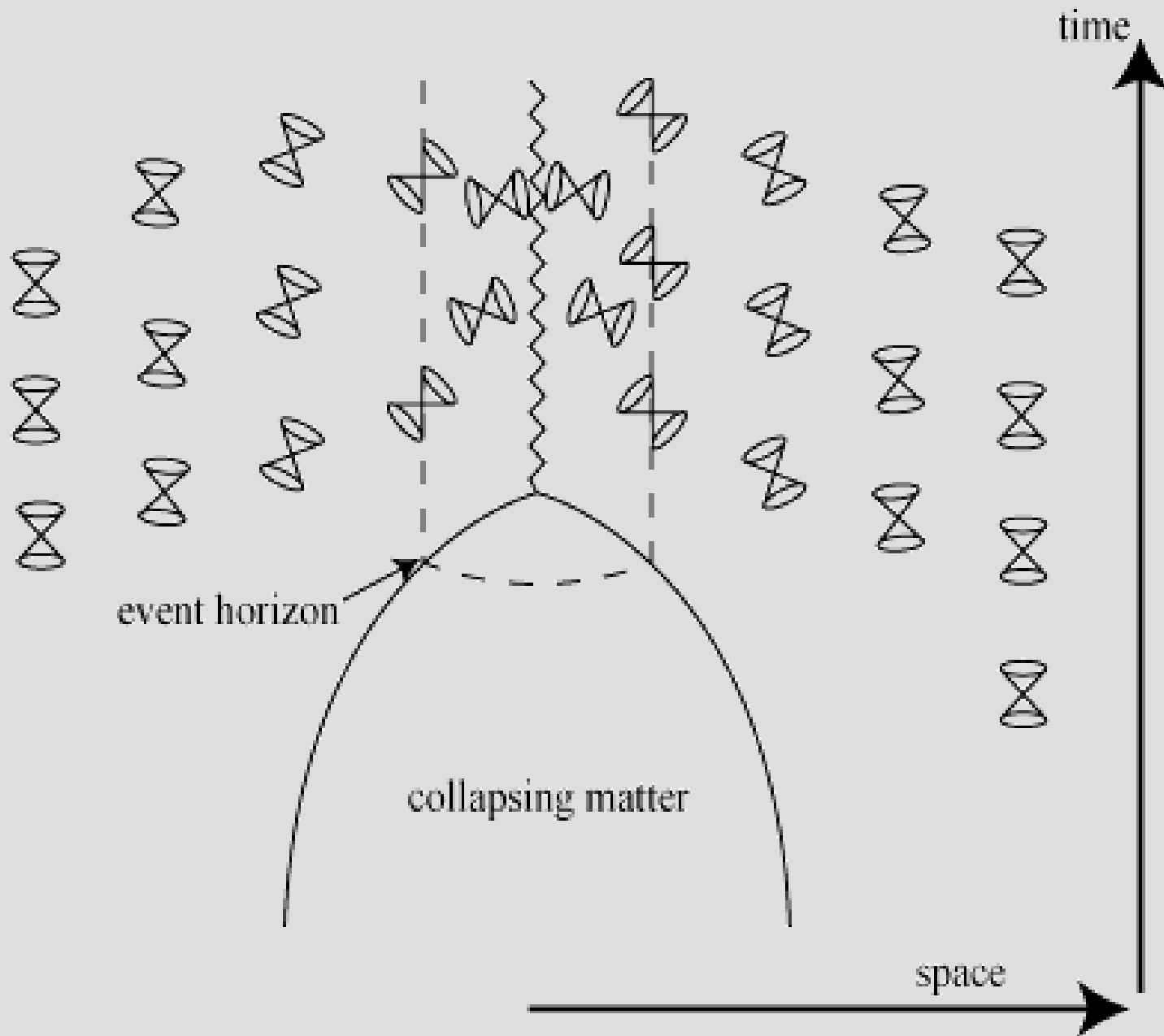
II. Event Horizon Telescope

III. The Image

IV. The Future of Black Holes

# Black Holes





The event horizon is "a perfect unidirectional membrane: causal influences can cross it in only one direction". —D. Finkelstein 1958



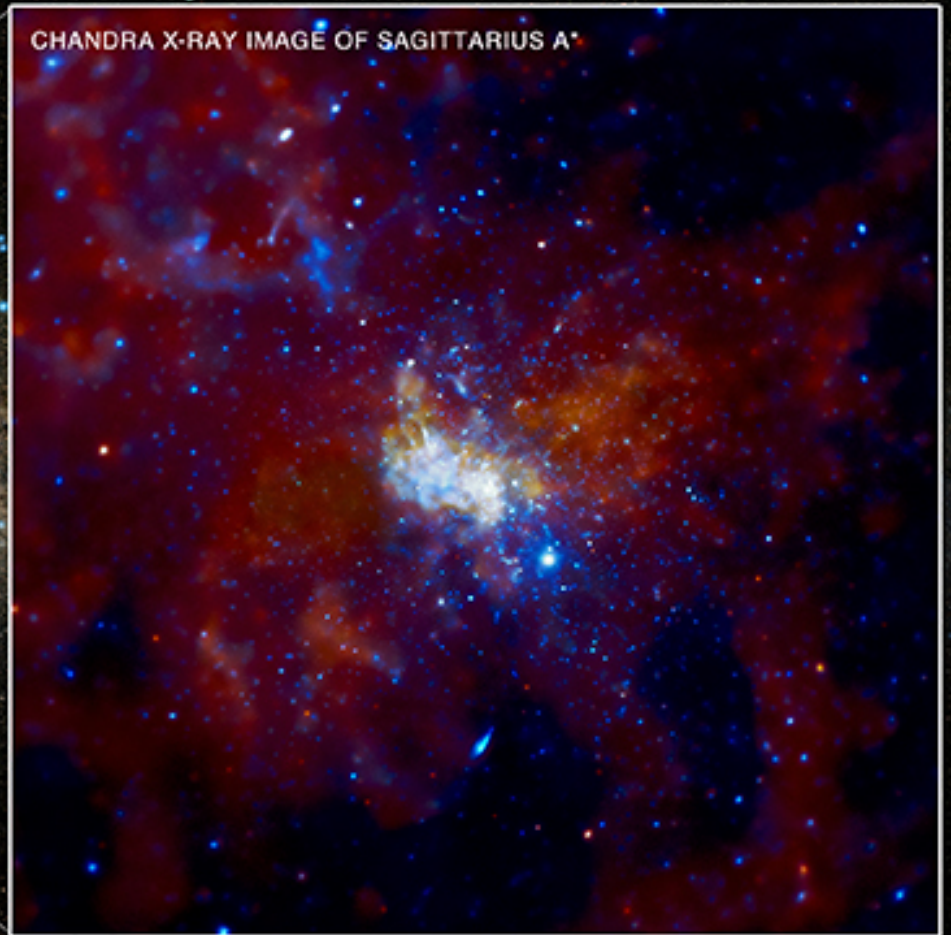




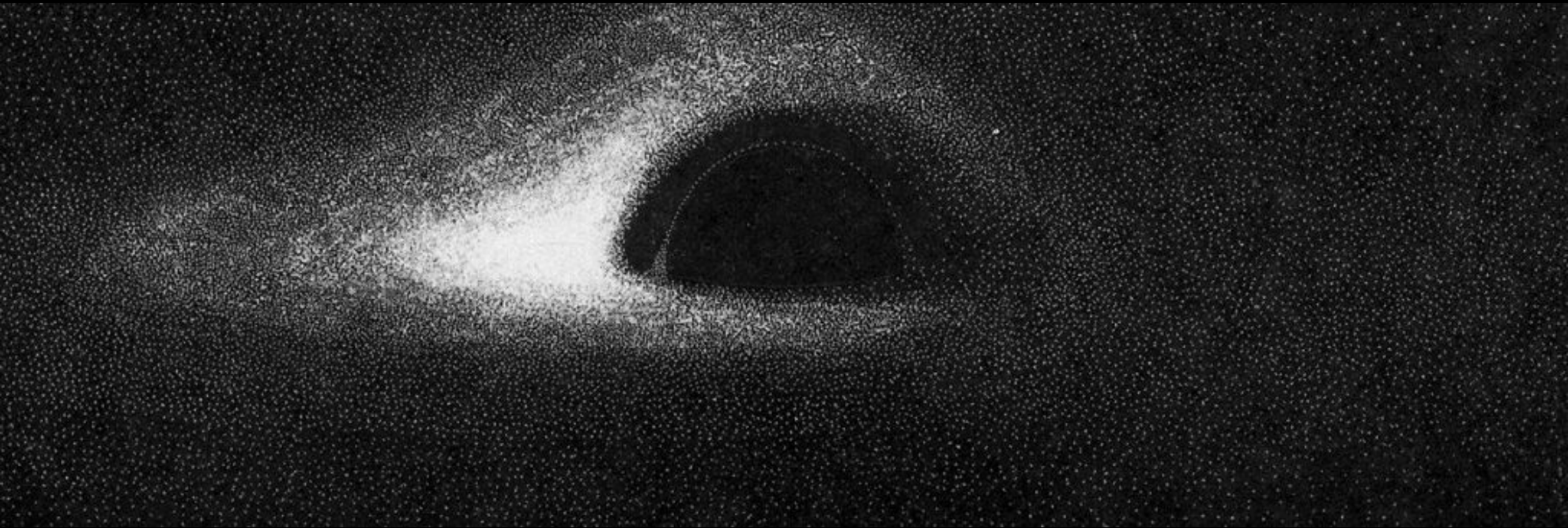
OPTICAL IMAGE OF THE MILKY WAY GALAXY BY E. SLAWIK



CHANDRA X-RAY IMAGE OF SAGITTARIUS A\*



In the 1970's Jim Bardeen, Jean-Pierre Luminet and others make first theoretical images of a Black Hole



# Imaging Black Holes: a Much Worse Problem

We characterize the size of a black hole with the radius of its event horizon, known as the Schwarzschild radius:

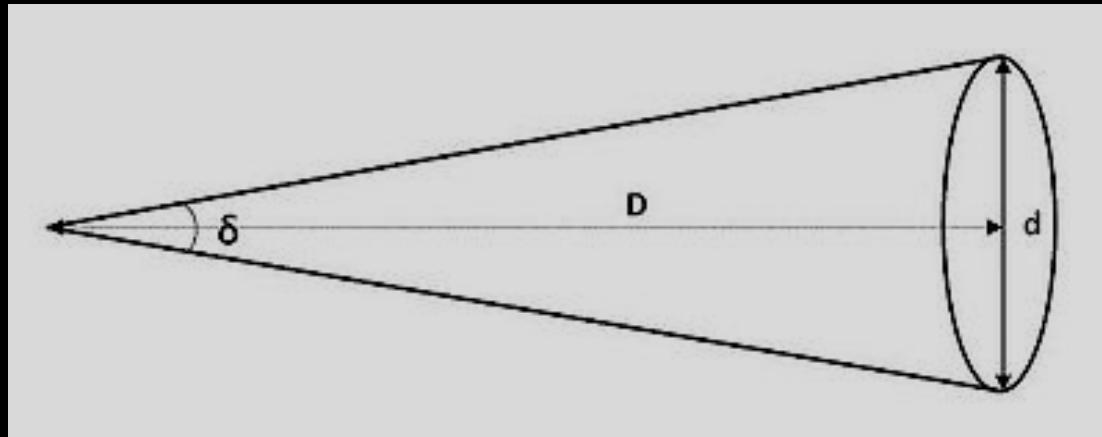
$$R_S = \frac{2GM}{c^2}$$

BH Sun	$R_S = 3 \text{ km}$
--------	----------------------

Sag A <sup>*</sup>	$R_S = 12 \times 10^6 \text{ km}$
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M87 <sup>*</sup>	$R_S = 19.5 \times 10^9 \text{ km}$
------------------	-------------------------------------

# Imaging Black Holes: a Much Worse Problem



Sun  $\delta = 0.53^\circ = 1913 \text{ asecs}$

BH Sun  $\delta = 8251 \mu\text{asecs}$

Sag A\*  $\delta = 20 \mu\text{asecs}$

M87\*  $\delta = 16 \mu\text{asecs}$

# Outline

I. Black Holes

II. Event Horizon Telescope

III. The Image

IV. The Future of Black Holes



# Event Horizon Telescope (EHT)

A Global Network of Radio Telescopes

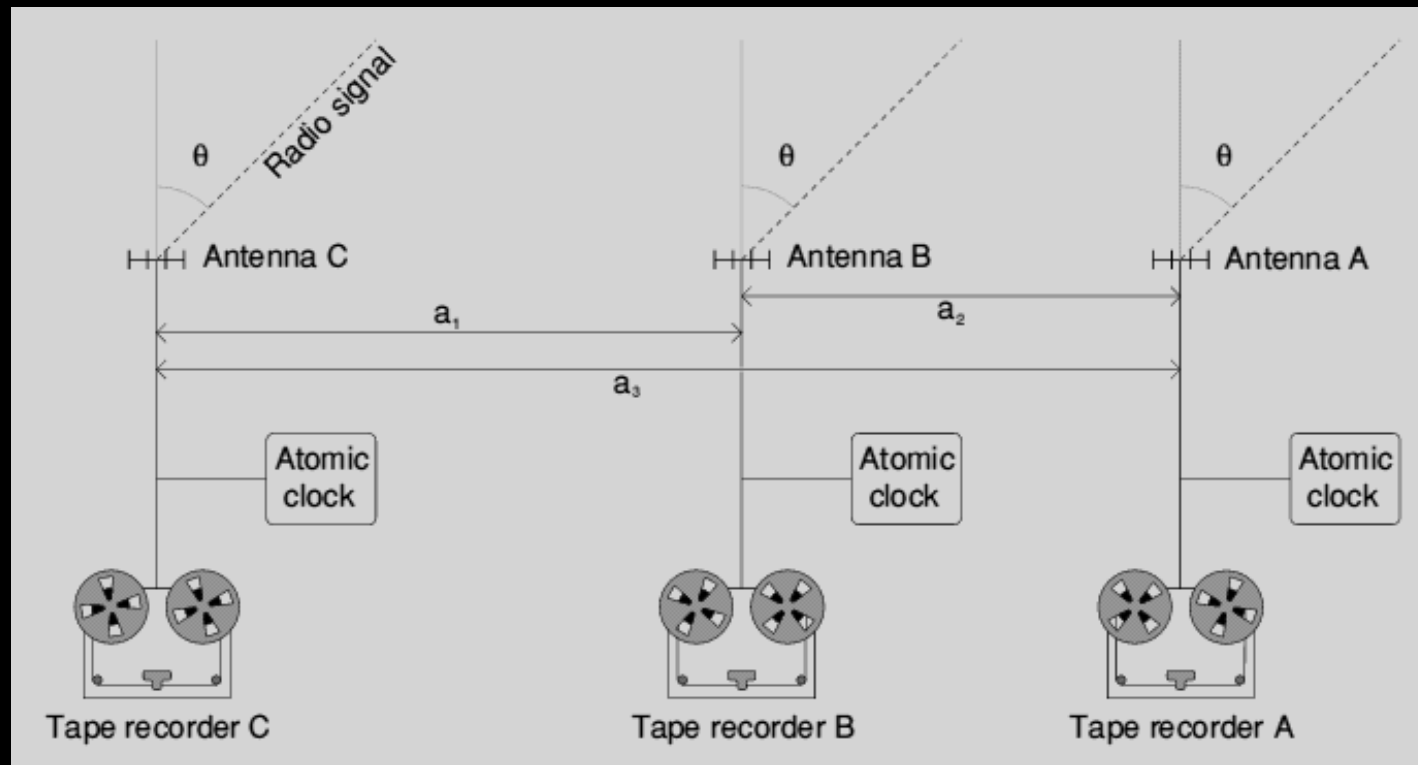
## 2018 Observatories



ALMA		Atacama Large Millimeter/ submillimeter Array CHAJNANTOR PLATEAU, CHILE
APEX		Atacama Pathfinder EXperiment CHAJNANTOR PLATEAU, CHILE
30-M		IRAM 30-M Telescope PICO VELETA, SPAIN
JCMT		James Clerk Maxwell Telescope MAUNAKEA, HAWAII
LMT		Large Millimeter Telescope SIERRA NEGRA, MEXICO
SMA		Submillimeter Array MAUNAKEA, HAWAII
SMT		Submillimeter Telescope MOUNT GRAHAM, ARIZONA
SPT		South Pole Telescope SOUTH POLE STATION
GLT		The Greenland Telescope THULE AIR BASE, GREENLAND, DENMARK
Kitt Peak		Kitt Peak 12-meter Telescope KIT PEAK, ARIZONA, USA
NOEMA		NOEMA Observatory PLATEAU DE BURE, FRANCE

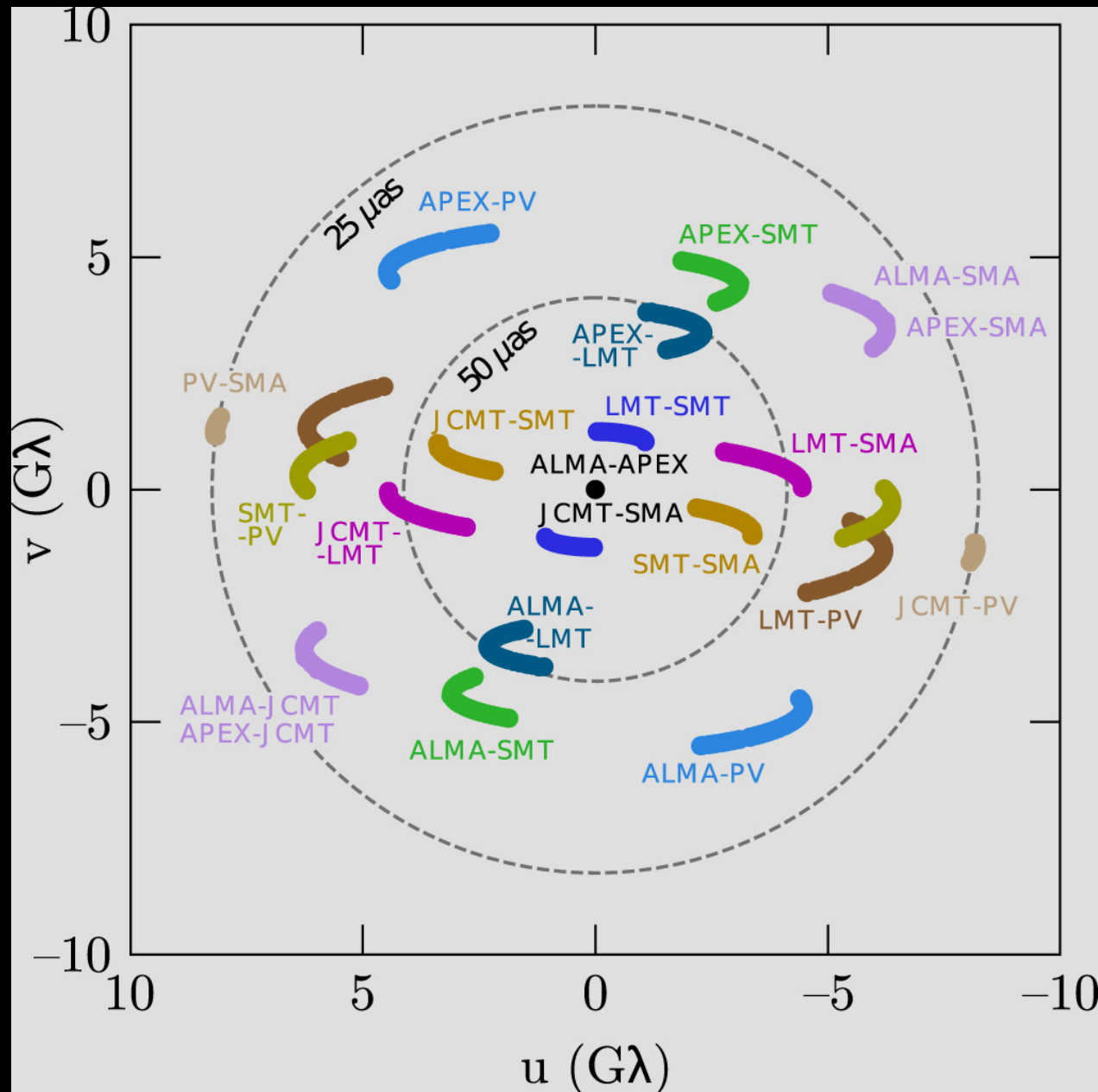
Observing  
in 2020

Signals from the telescopes are recorded locally and time stamped using extremely precise and stable atomic clocks



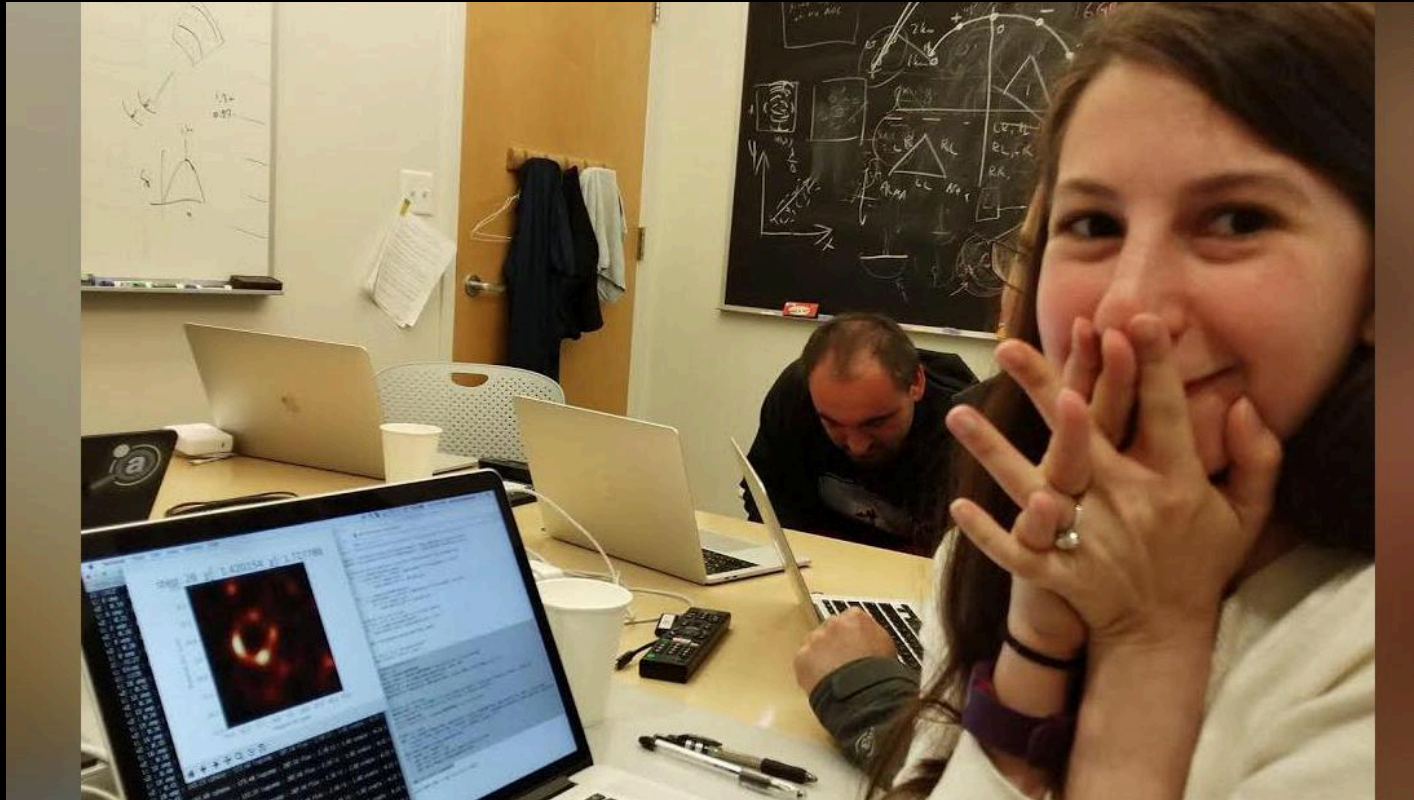
The data are collected and compared carefully; they use both time stamp and predicted time delays to align the signals

# Combined coverage over the 4 nights of observing





The collaboration split into 4 teams that each independently reconstructed images from the interferometry data



[www.ted.com/talks/  
katie\\_bouman\\_what\\_does\\_a\\_black\\_hole\\_look\\_like?  
language=en](https://www.ted.com/talks/katie_bouman_what_does_a_black_hole_look_like?language=en)

# Outline

I. Black Holes

II. Event Horizon Telescope

III. The Image

IV. The Future of Black Holes





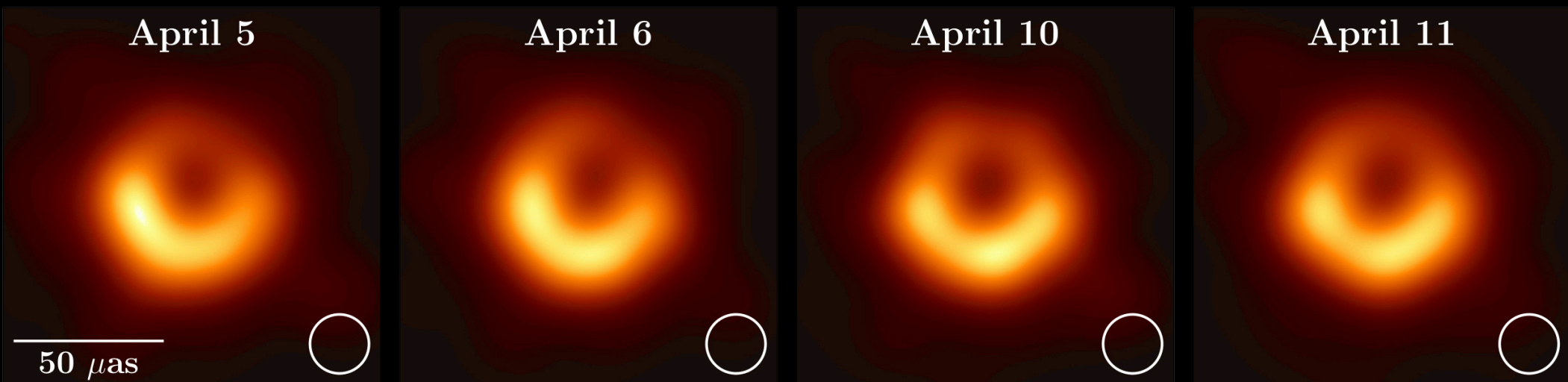
April 5

April 6

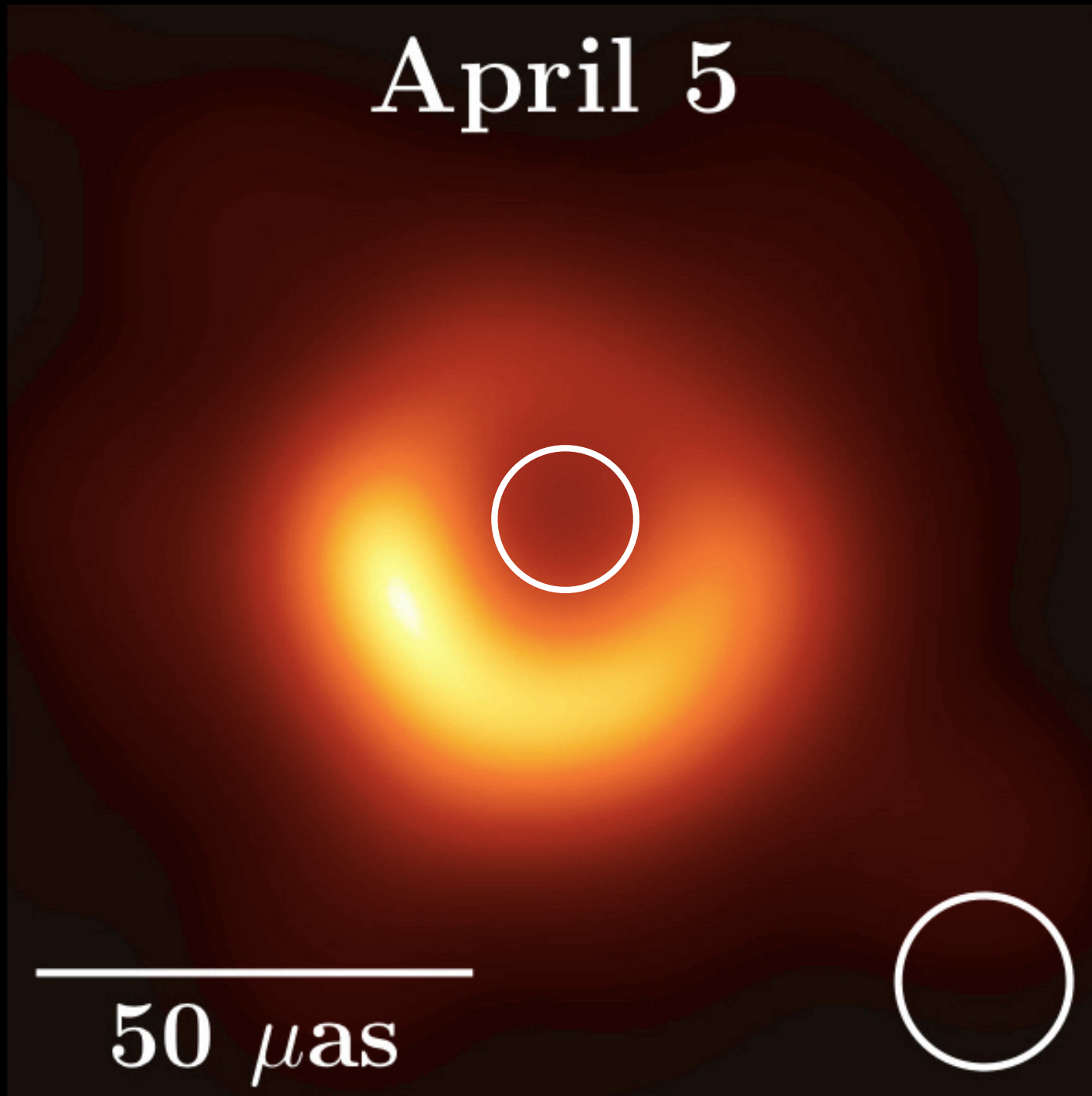
April 10

April 11

50  $\mu\text{as}$

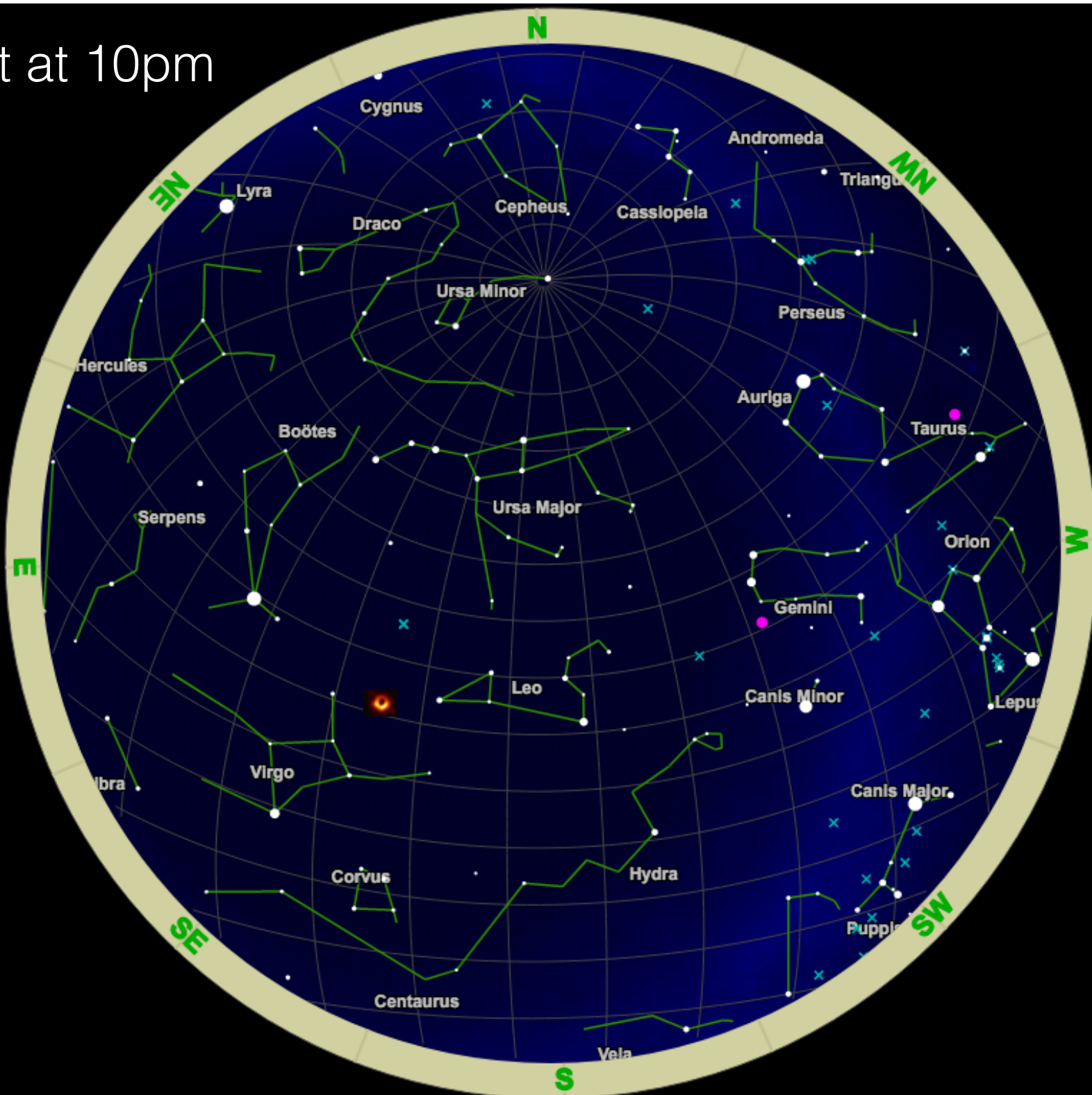


April 5



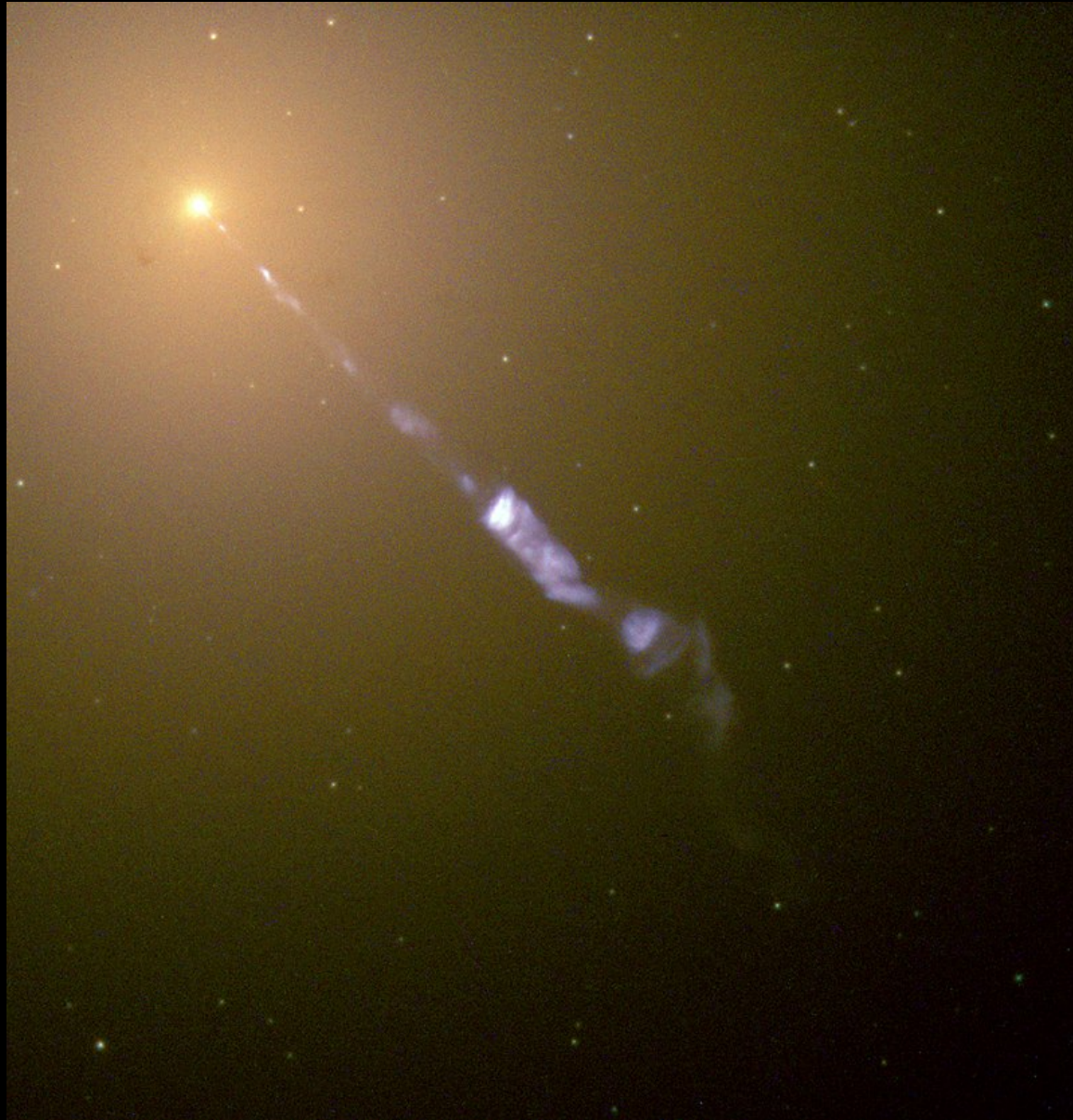
$50 \mu\text{as}$

Tonight at 10pm



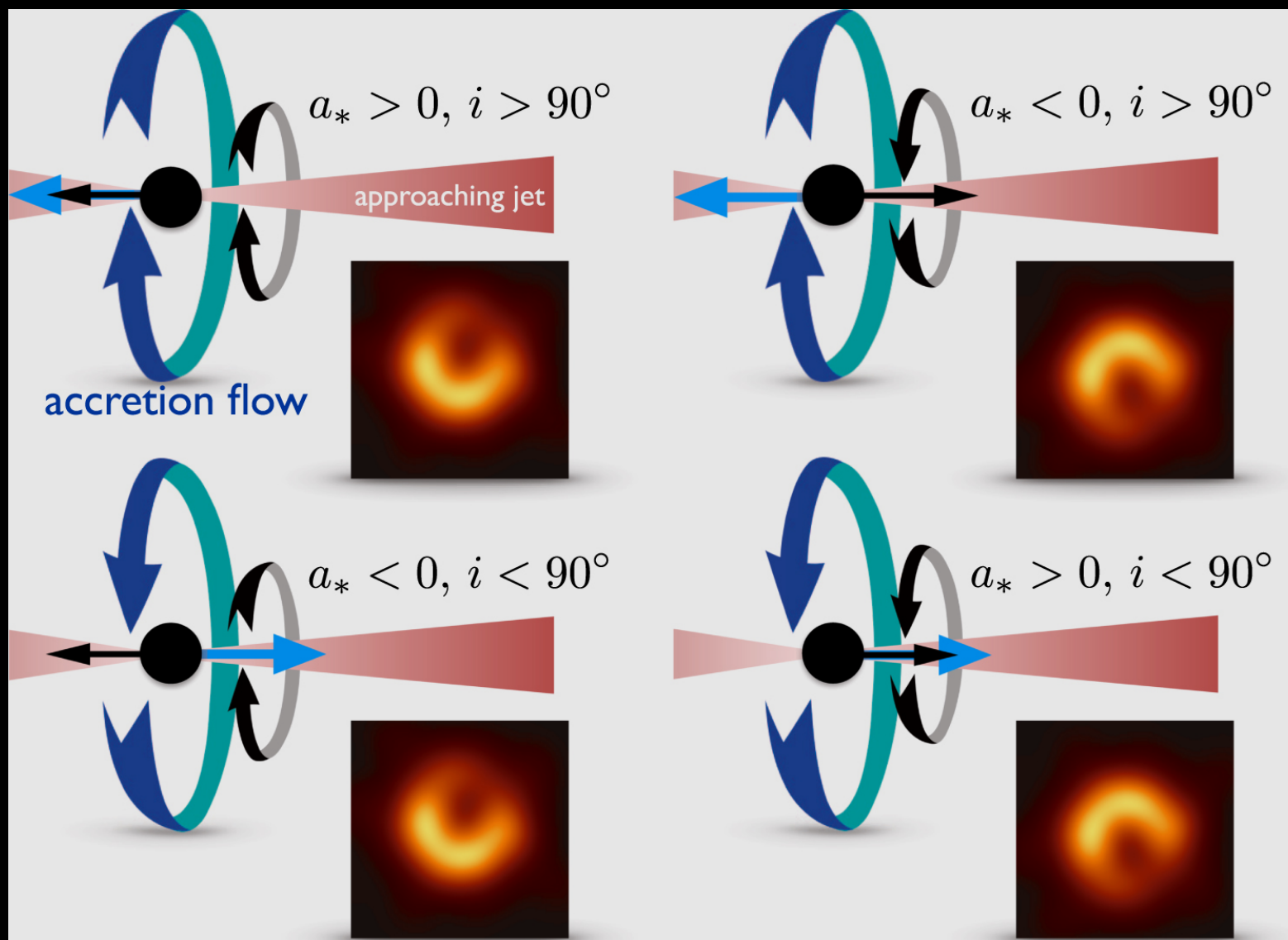


The image is consistent with a spinning black hole





These images are consistent with a spinning black hole.  
How is it spinning?



Why is the image blurry?

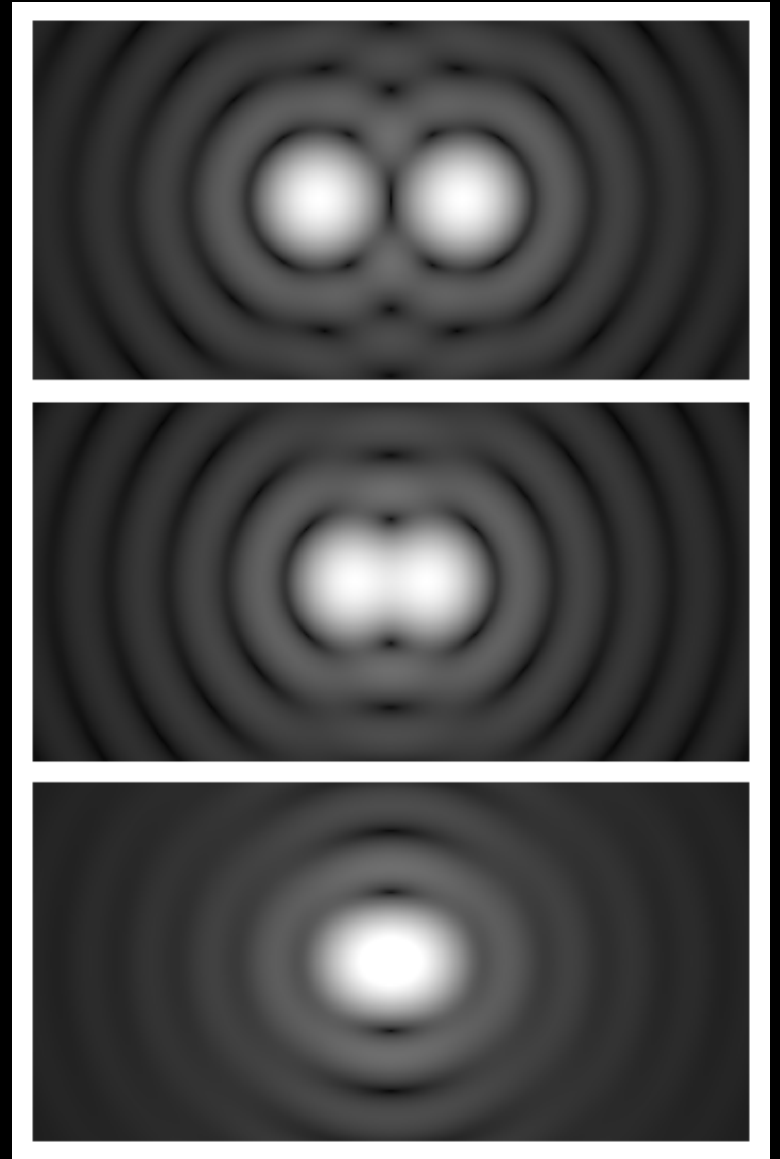


$$\theta \sim \frac{\lambda}{D}$$

$$\lambda = 1.3 \text{ mm}$$

$$D = 2R_E = 12741 \text{ km}$$

$$\Rightarrow \theta = 21 \mu\text{asecs}$$



# Outline

I. Black Holes

II. Event Horizon Telescope

III. The Image

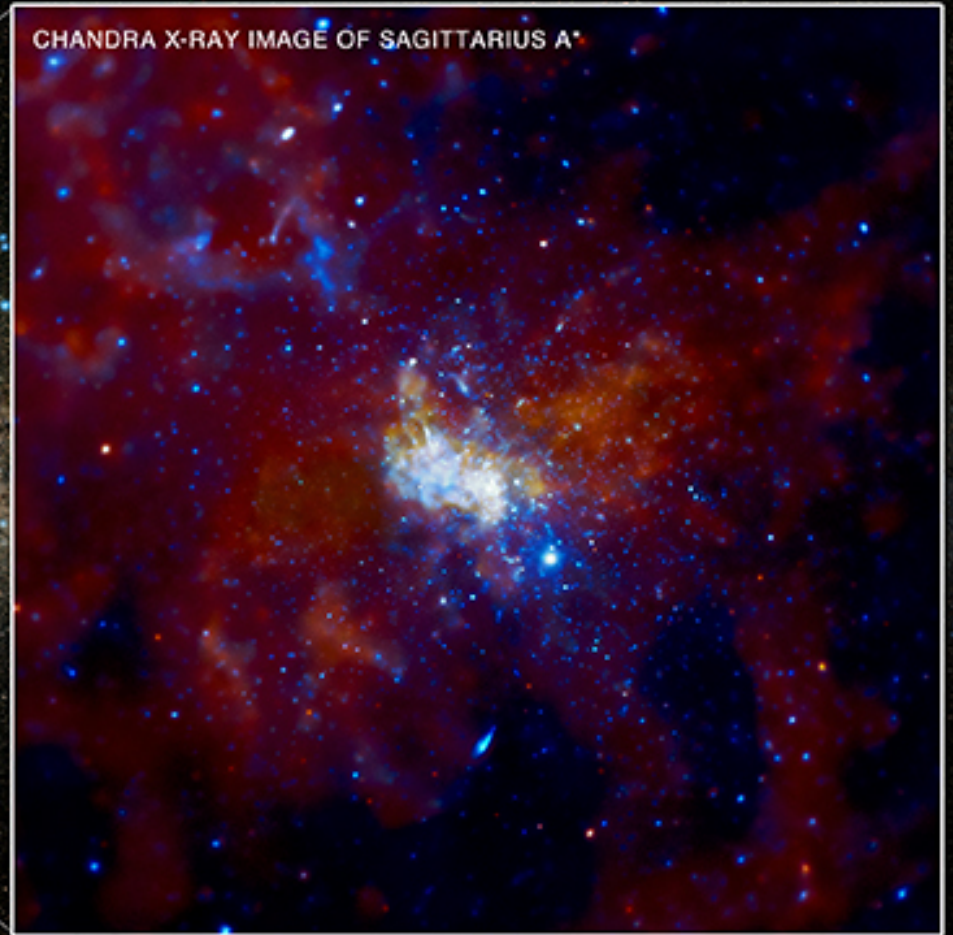
IV. The Future of Black Holes

We are waiting for images of our own black hole Sag A\*

OPTICAL IMAGE OF THE MILKY WAY GALAXY BY E. SLAWIK

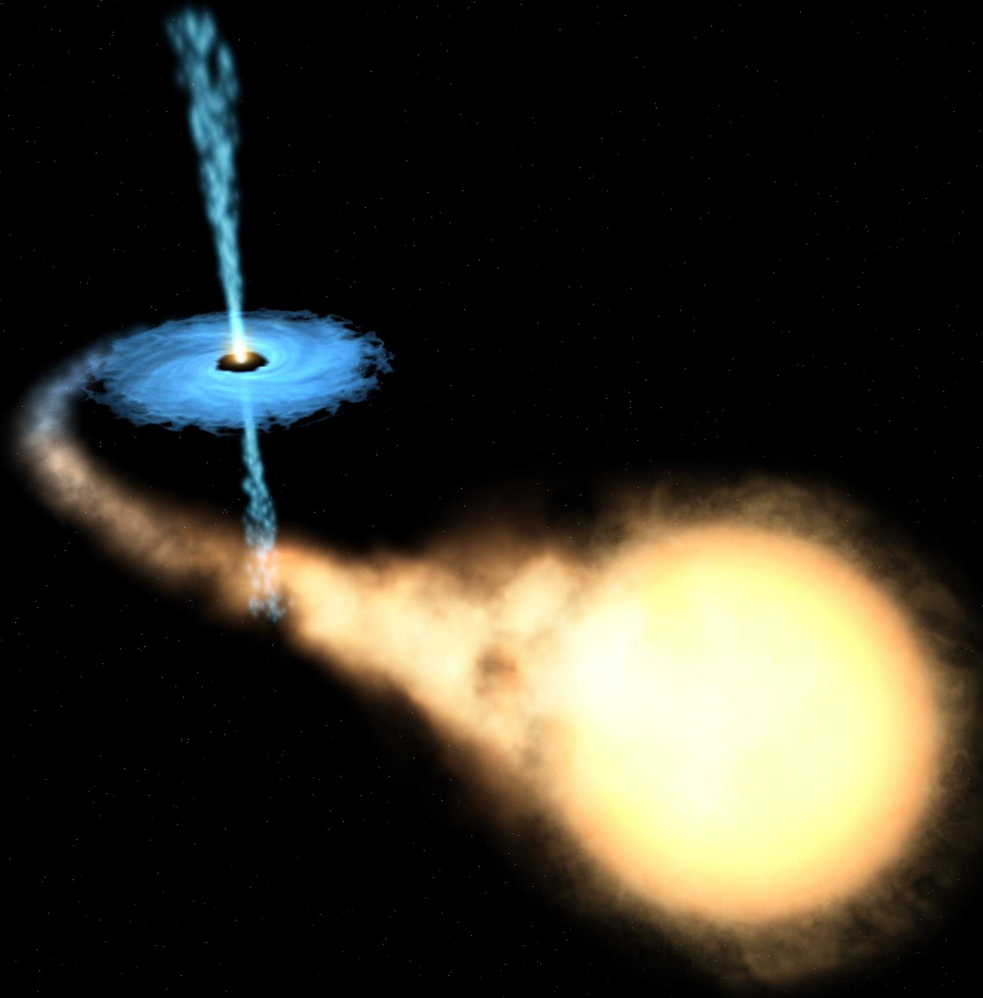


CHANDRA X-RAY IMAGE OF SAGITTARIUS A\*

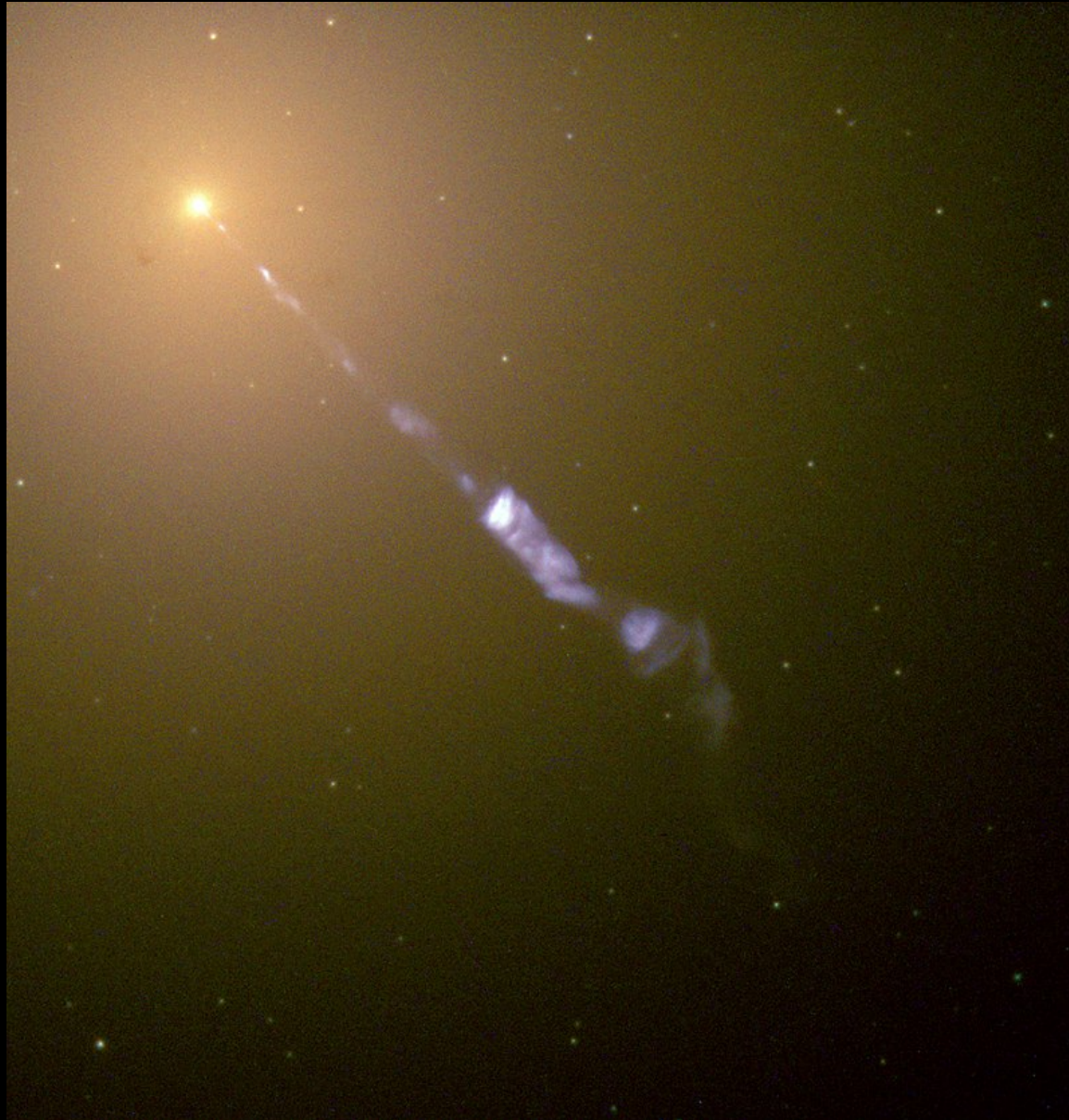




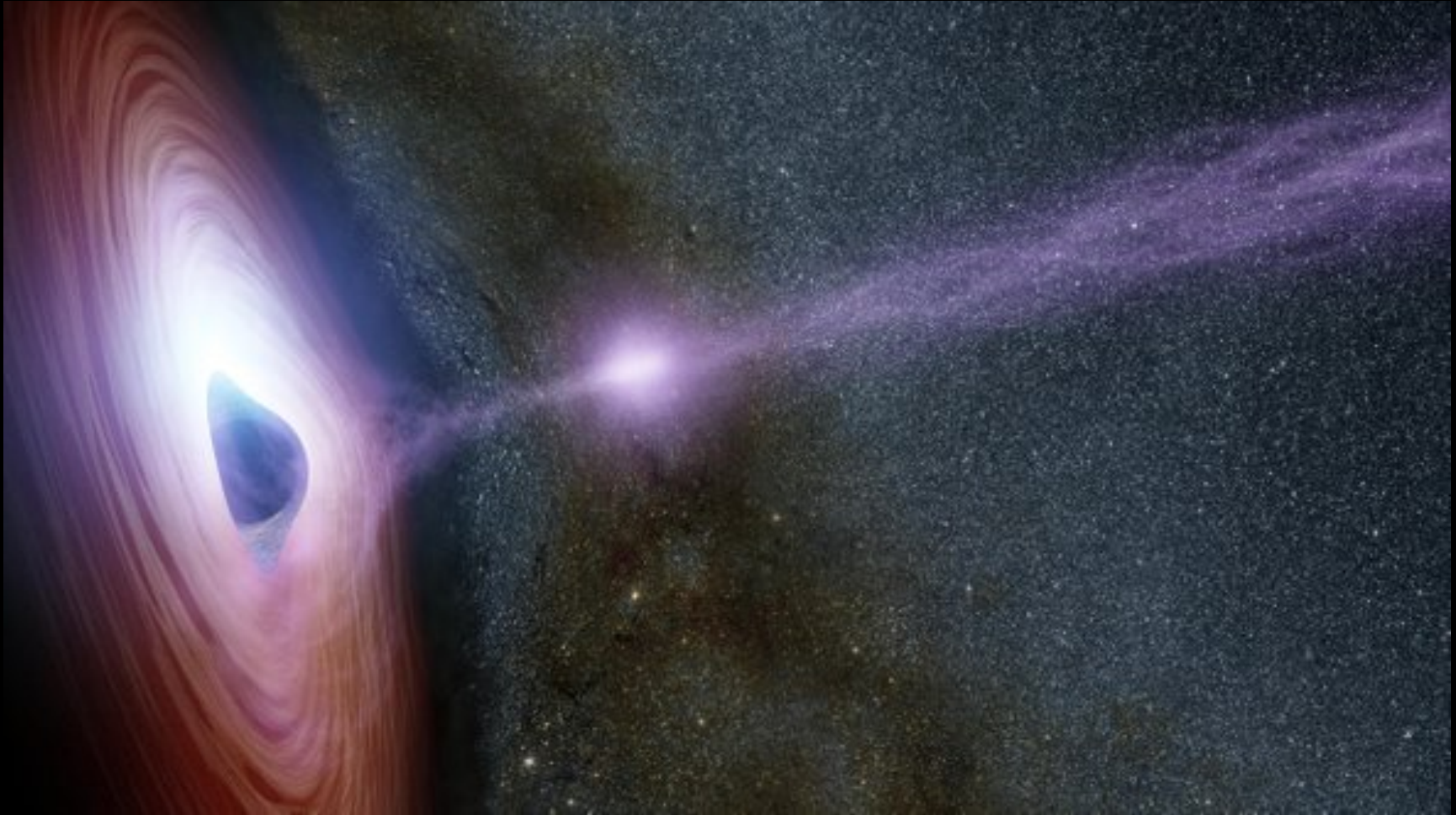
We want to see a movie of black holes eating



We want to know how supermassive black holes drive jets?



Are there quantum effects near the event horizon?



In the last few years the door has opened to study black holes throughout the skies and at a variety of length scales



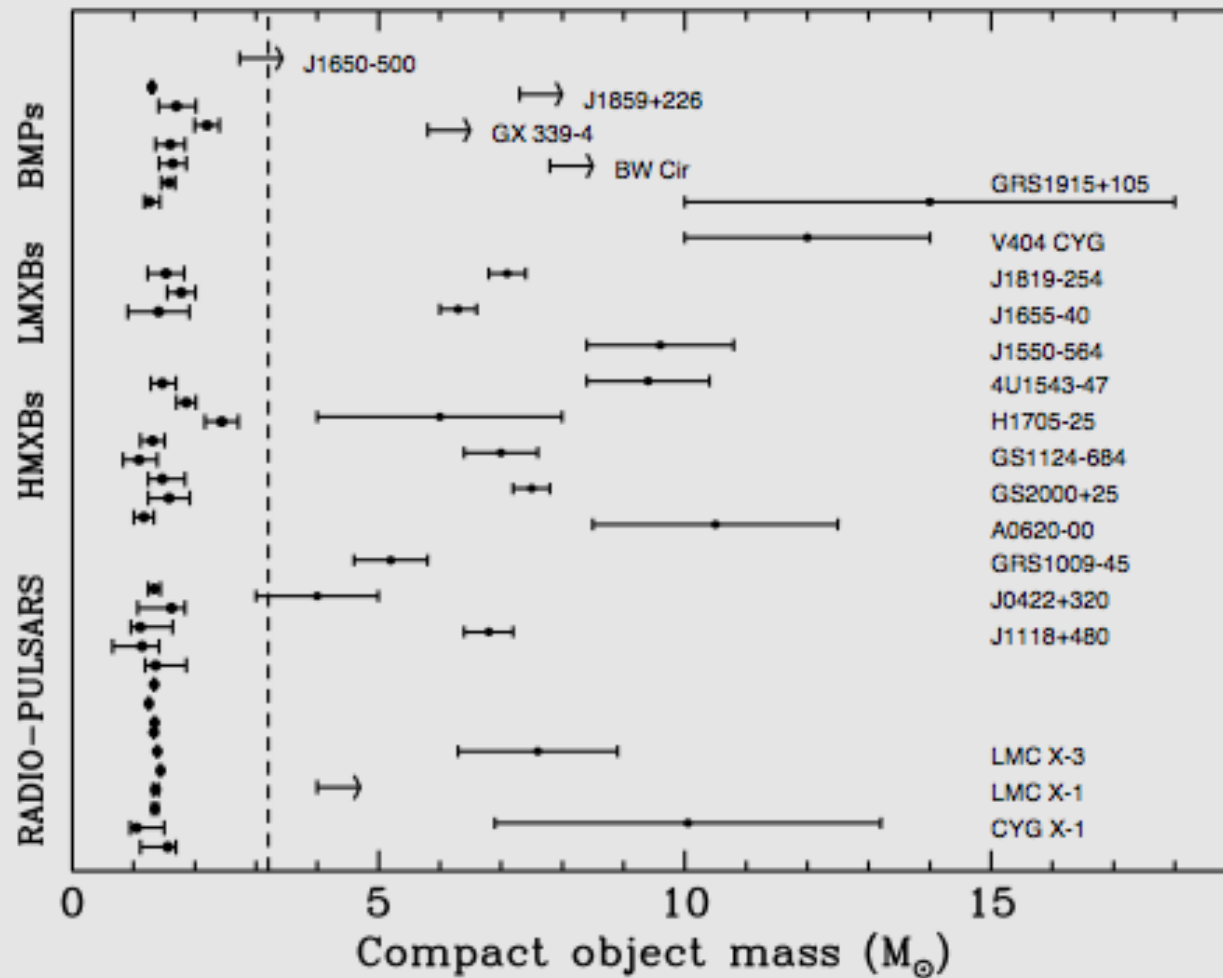
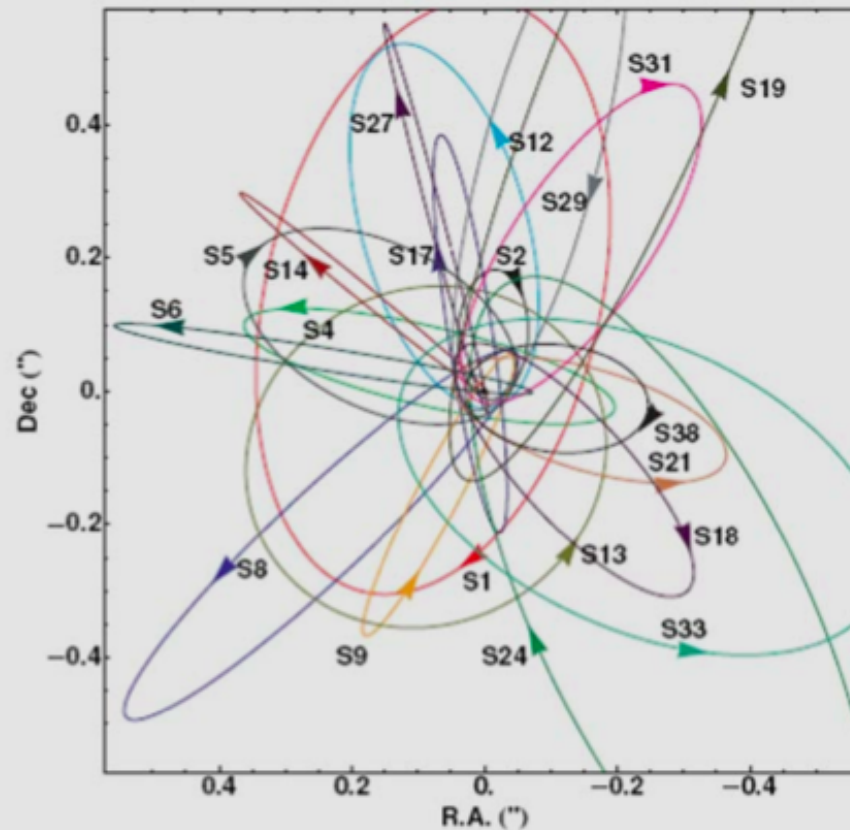
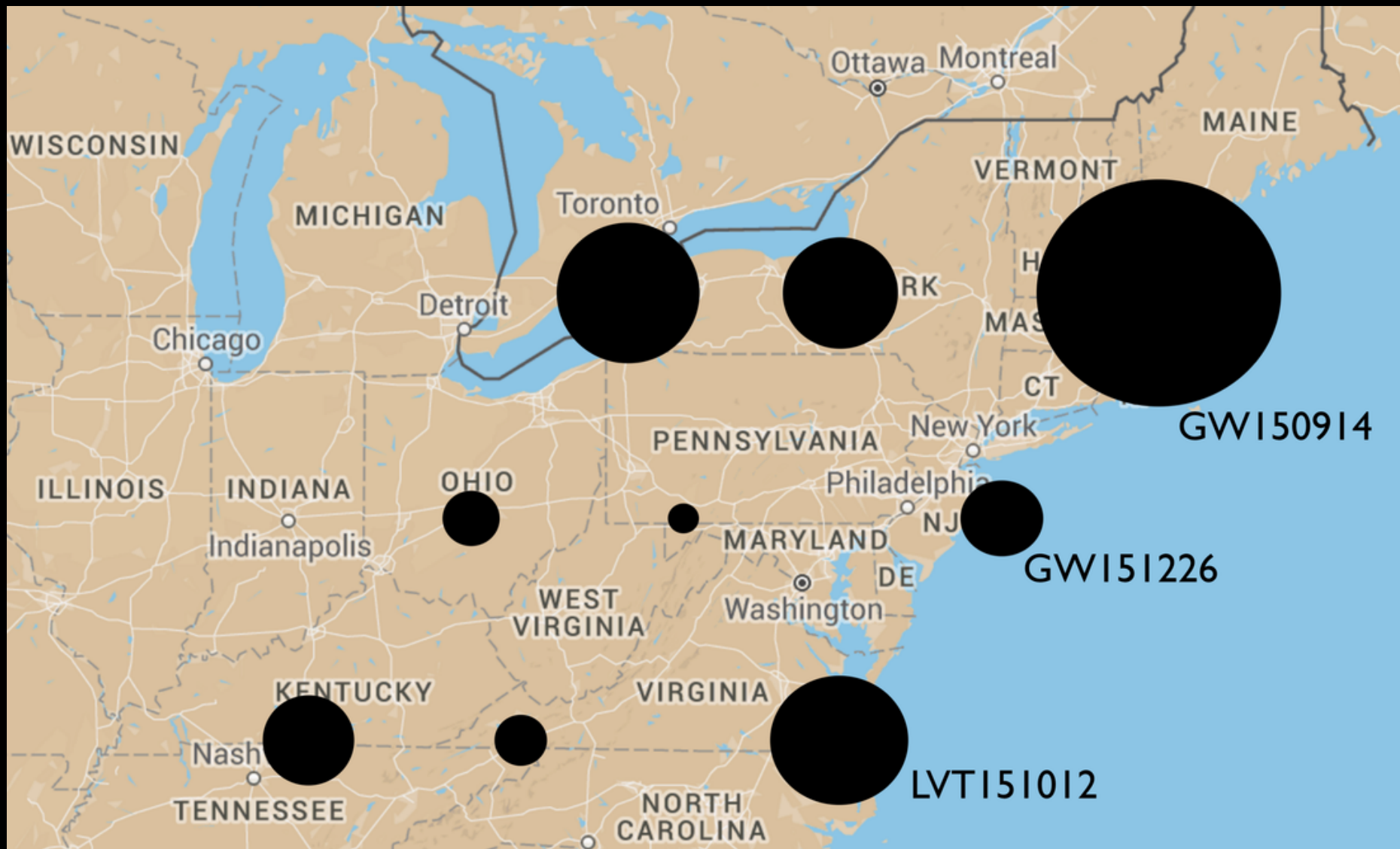


Figure 6. Mass distribution of compact objects in X-ray binaries. Arrows indicate lower limits to BH masses. Figure reproduced from Casares (2007).



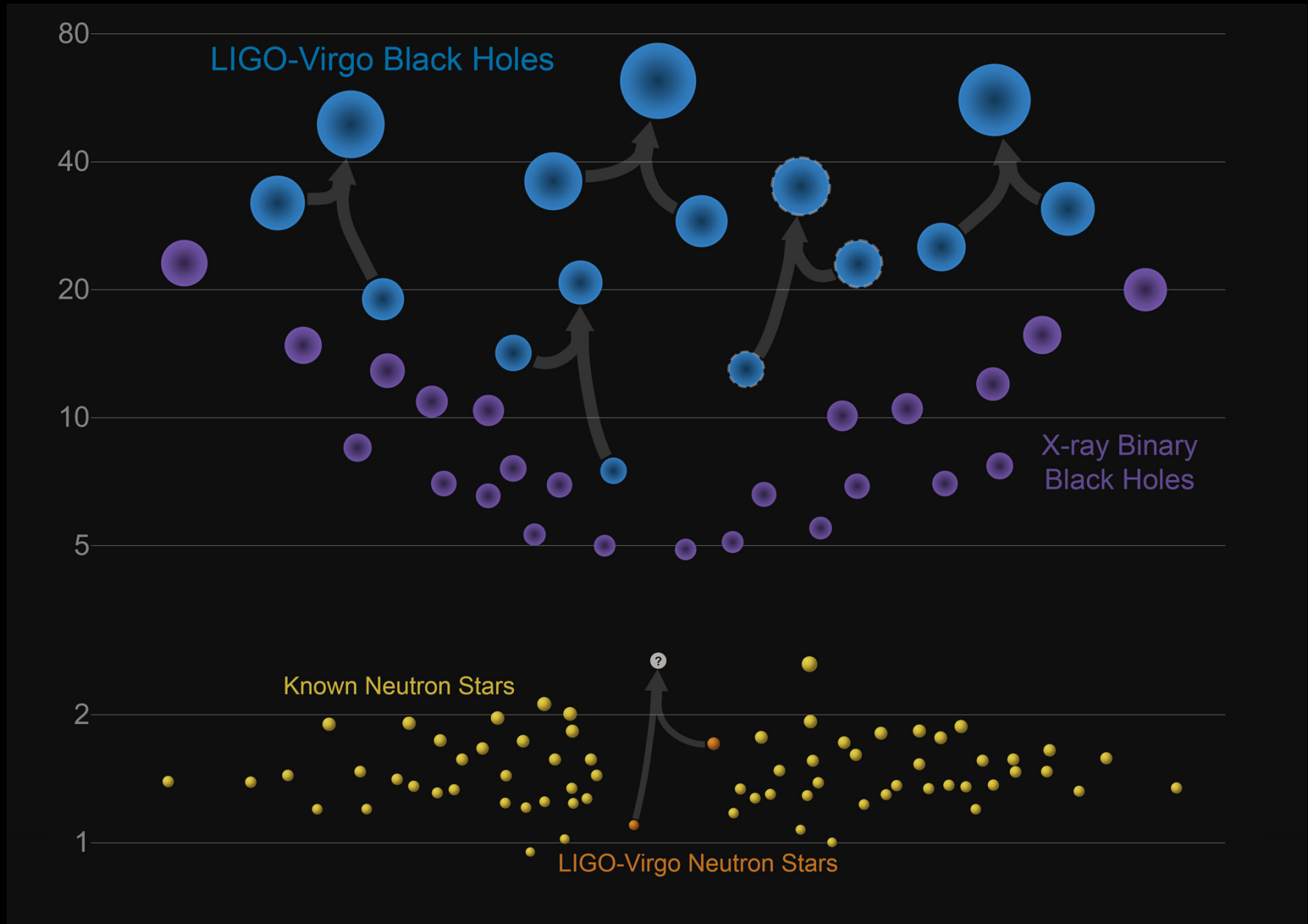
**Figure 4.3.2.** A summary of 20 of the  $\sim 30$  S-star orbits delineated by the most recent orbital analysis of Gillessen et al. (2009b) <sup>5</sup>.

Tracking orbiting stars at the center of the Milky Way accurately established the mass of Sag A\* at  $4.3 \times 10^6 M_{\odot}$ .





# In solar masses



They offer a remarkable laboratory for thermal physics, strong gravity and one day, I hope, for quantum gravity

The End