Today's Outline I hast Lecture: Completed discussion of Normal modes I Introduction to Chaos le Nortiveux Mechanics

Lec-ture 30 
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\pm
$$
 Last lecture  $PV_{H}$   
\nNovember 9<sup>th</sup>, 2011  $F$  and normal mode  $ECM$ .  
\n $\vec{M} \cdot \vec{g} = -\vec{k} \cdot \vec{g}$   
\n $\vec{M} \cdot \vec{g} = -\vec{k} \cdot \vec{g}$   
\n $\vec{M} \cdot \vec{h} = A_{ik}(\vec{0}) = \sum_{d} M_{d} \left(\frac{\partial \vec{f}_{d}}{\partial g_{i}}\right) \cdot \left(\frac{\partial \vec{f}_{d}}{\partial g_{j}}\right)$   
\n $Evaluate \vec{g} = \vec{g}$ 

$$
K_{jk} = \frac{\partial^2 U}{\partial g_j \partial g_k} \Big|_{g=0}
$$

Notice that you, in fact, ne clu't write down the Lagrangian at all<br>fore these problems; you can just find  $\widetilde{M}$  and  $\widetilde{K}$  and proceed as before.<br>Guess:  $\vec{g}(t) = \overline{k}e^{\vec{z}(t)}, \overline{\vec{z}}(t) = \vec{a}e^{i\omega t}$ 

and<br>Solve:  $\begin{cases} det(\vec{k} - \omega^2 \vec{M}) = 0 \text{Normal Freg.} \\ (\vec{k} - \omega^2 \vec{M}) \vec{a} = 0 \text{Normal nodes} \end{cases}$ 

Your text goes through another<br>example :<br>Check it out! II chaos & Noutineau Mechanics A system is said to be chaotic if it obeys deterministic E.O.M. but still the motion cannot be predicted on long time scales.

Most realistic systems are nonlinear:  
\n
$$
eq_{1}
$$
,  $m\frac{1}{12} = \frac{GmM\hat{r}}{r^2}$  or  $\frac{d^2x}{dt^2} + \mu(x^2-1) \frac{dx}{dt} + x = 0$   
\nOF 4 $en_{1}$ , but not always nonlinear  
\nSystems, exhibit chaos.  
\nNonlinearly changes many things but  
\nimportant amongst these is that nonlinear  
\nrespectants do not allow auper position  
\nSystems do not allow auper position  
\nD  $(\alpha_1 x_1H) + \alpha_2 x_2H) = \alpha_1 D x_1 + \alpha_2 D x_2$  [linear]

The E.O.M is: a ssumed<br>damping  $\mathcal{I}\ddot{\phi} = \Gamma$ =>  $mL^{2}\phi = -mgLsin\phi - b\overline{v}L + LF(t)$ = -mgL  $sin\phi - bL^2\phi + LF_0 \cos(\omega t)$ Assume Sinusoidal drive at  $\Rightarrow$   $\phi + \frac{b}{a}\phi + \frac{a}{b}\sin\phi = \frac{r}{aL}\cos(\omega t)$ Let  $2\beta = \frac{b}{n}$ ,  $\beta$  the "damping constant"

$$
D(a_{1}x_{1}+a_{2}x_{2})\neq a_{1}Dx_{1}+a_{2}Dx_{2}
$$
  
\n $[D$  nonlineer]  
\nWe'll explore all of this  
\nwith a particular nonlinear  
\nsystem: the damped driven  
\npenduhan (DPP)  
\n $\frac{1}{P}$  Fth)  
\n $hr$ 

 $\omega_0 = 13/L$  the "natural frequency" and  $\gamma = \frac{F_{o}}{ML\omega_{o}^{2}}$  the "drive strength" d'is dimensionless (by design) Since F. /ML has dimension (time)<sup>2</sup><br>The name names sense because  $Y = \frac{F_e}{mL\omega_o^2} = \frac{F_e}{mg}$ for  $Y \nleq 1$ ,  $F_o$  any and has<br>a small effect, while for<br> $Y \ngeq 1$ ,  $F_o$  > any and has a large effect.

Finally then,  $\dot{\phi}$  + 2p $\dot{\phi}$  +  $\omega_0^2$  sind = 8  $\omega_0^2$  cosut To appreciate how wild chaos is We first need to revind ourselves of what we would naively expect. For small oscillations.  $\dot{\phi} + 2 \rho \dot{\phi} + \omega_{0}^{2} \phi = \gamma \omega_{0}^{2} \cos \omega t$ we get a damped driven oscillator! If we now increase the drive strength the amplitude will increase and More<br>terms will natter! Six  $\phi \approx \phi - \frac{1}{6} \phi^3 + \cdots$  $\phi$  +2p  $\phi$  +wo  $(\phi - \frac{1}{6} \phi^3) = \gamma \omega_0^2$  cosut This is a small perturbation, so quess  $\phi$  (A)  $\approx$  A cos(wt-8) When we put this guess into the E.O.M. we get a term  $\phi^3 = A^3 \cos^3(\omega t^{-5})$ . A useful try identity is

We expect that for weak driving P3/4 and small uscillations: The initial behavior of the pendulum depends on the initial conditions, but this transient motion dies out rapidly, and the protion approaches a unique attractor! j'aine frequency.<br>p(t) = Acos (wt-s) See plot (Fig 1).

 $cos^3 x = \frac{1}{4} (cos 3x + 3 cos x)$ Then a simplification of the  $\phi^3$ tern will include cos(43[ut-8]) but the drive has a prescribed frequency is and so this tripled frequency must be concelled by One of  $\phi$ ,  $\phi$ ,  $\phi$  (in fact all three), and a better guess is  $\phi$  (t) = A  $cos$  (wt-s) + B  $cos$  3 (wt-s)

As the drive increases, this pattern  
confluues and we get higher harmonics  
(multiples & n of (w+-s)). This  
Slady nolifies the attractor solution  
(see Fig.2)  
as we approach 
$$
Y=1
$$
. Near  $Y=1$   
we get dramatically different results!  
III Miftern 2 Histogram &  
Results.

Now plan for the final exan 
$$
\frac{pq}{l}
$$



Figure: 1) Taken from Taylor p464



Figure: 2) Taken from Taylor p466