

## Today's Outline

I Prelude

II Interlude: Lagrangian  
& First order?

0. Announcements: (i) For this material I depart from Taylor. (ii) Last HW due last day of class  
(iii) Brunch at my house Dec 4<sup>th</sup>, 10:30 am

This prelude begins to answer this question.

Consider a harmonic oscillator; we usually write the E.O.M. as

$$m\ddot{q} = -kq \quad \text{or} \quad \ddot{q} = -\omega_0^2 q,$$

a second order, ordinary differential equation. Equivalently we could write this as two first order ordinary differential equations: Define a

## Lecture 33

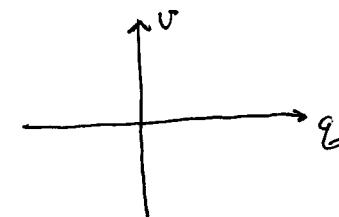
November 18<sup>th</sup>, 2011

## I Prelude

PL/4

Hamiltonian mechanics is a third approach, distinct from both the Lagrangian and Newtonian approaches we have been studying so hard. A question suggests itself: Why consider another approach?

new coordinate  $v$ , the velocity. With these two coordinates  $(q, v)$  the arena for our dynamics becomes



and we call it "velocity phase space" or "state space". The equations of motion are

$$\dot{g} \equiv \frac{dg}{dt} = v$$

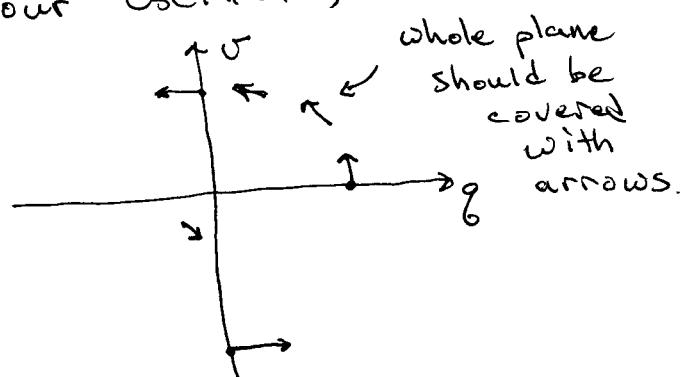
$$\dot{v} = -\omega_0^2 g$$

As far as solving the E.O.M goes this does not represent major progress. However, there are technical and geometrical reasons that it is advantageous. Let's begin with geometry: In this form the E.O.M. represent a vector field,

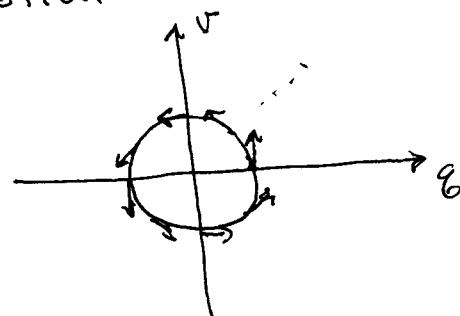
An example of a vector field, familiar to many of you, is the electric field:  $\vec{E}(x, y, t)$ . Instead of drawing all of the little arrows we usually visualize the field lines. The vector field is tangent to the field lines at every point of space, and in this context we usually call the field lines "integral curves" of

that is, a set of equations  $\frac{dx_i}{dt} = v_i$  that tell us how our coords change (in time) at every point of our velocity phase space.

For our oscillator,



the vector field. Similarly our first order equations have integral curves and these are precisely the orbits of the motion



This begins to introduce the geometry.

Now for some technical results:

Under appropriate mathematical conditions,

(e.g. the R.H.Ss are continuous and differentiable)

Solutions

- Exist and
- Are unique for given initial conditions  
 $\vec{q}_0, \vec{v}_0$ .

These are things physicists often don't worry about, but it's nice to have a proof

(partially)

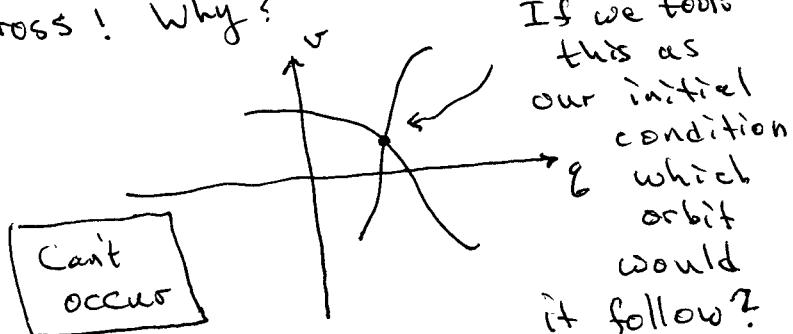
Hopefully you are convinced that there is much of interest when you write the E.O.M. in first order form.

## II Interlude

Writing the E.O.M. in first order form turns out to be clumsy in the Lagrangian formulation. Let's show this. To get E.O.M. in first order form we need

$$\ddot{\vec{q}} = \vec{\sigma} \quad \text{and} \quad \ddot{\vec{v}} = \ddot{\vec{g}}$$

here (see Arnold, Hirsch & Smale, or Hale). The second point has a nice geometrical interpretation: orbits in velocity phase space do not cross! Why?



the tricky part of this is the acceleration

$$\ddot{\vec{g}} = \ddot{\vec{g}}(\vec{q}, \vec{v}, t).$$

The acceleration is hiding in the Euler-Lagrange equation.

Let  $L = L(\vec{q}, \vec{v} = \dot{\vec{q}}, t)$  then, we can evaluate the  $\frac{d}{dt}$  term by the chain rule

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}_i} \right) = \frac{\partial L}{\partial q_i} \Rightarrow$$

$$\frac{\partial^2 L}{\partial \dot{q}_i \partial \dot{q}_j} \ddot{q}_i + \frac{\partial^2 L}{\partial \dot{q}_i \partial q_j} \dot{q}_j + \frac{\partial^2 L}{\partial \dot{q}_i \partial t} = \frac{\partial L}{\partial q_i}$$

P4/  
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formulation we will find  
first order E.O.M. expressed  
in a very simple form.

To isolate  $\ddot{q}_i$  we need to invert  
the matrix

$$M_{ij} = \frac{\partial^2 L}{\partial \dot{q}_i \partial \dot{q}_j}$$

and the resulting equation is  
complicated. In the Hamiltonian

Preview: Next week we will  
mathematically capture how  
Fig 1 encodes a curved line  
even though it was constructed  
solely by drawing straight lines:

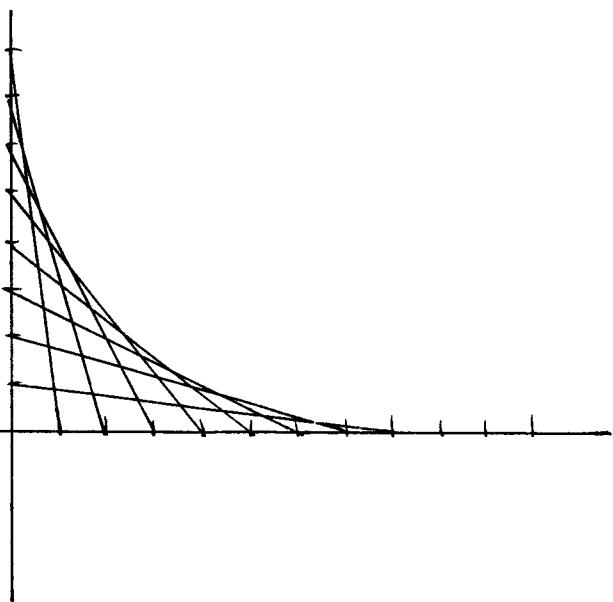


Fig.1: Asteroid exemplifying the Legendre Transform.