

Today's Outline

Lecture 3

I Geometry of Relativity ^{P1/5}

I. Geometry of Relativity

Jan. 24th, 2012

II. Recap relativistic effects
& transformations

III. Four Vectors

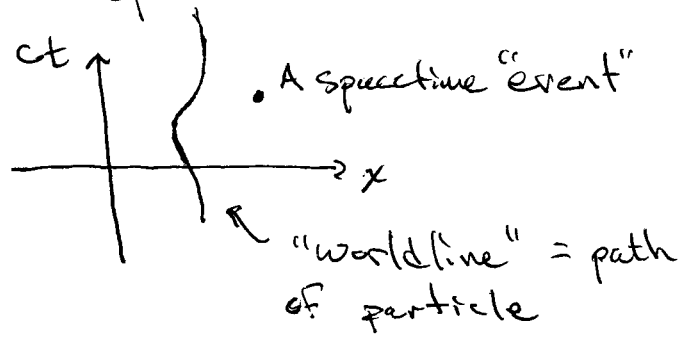
~~IV. S.R. Kinematics &
Dynamics~~

Last week we decided we would study ^{the} geometry of space and time. We decided to do this locally using the line element

ds^2 = squared distance
between nearby points

As you know, Einstein argued that we should study space and time together, that is spacetime.

~~We call this~~ Consider flat spacetime (or Minkowski space)



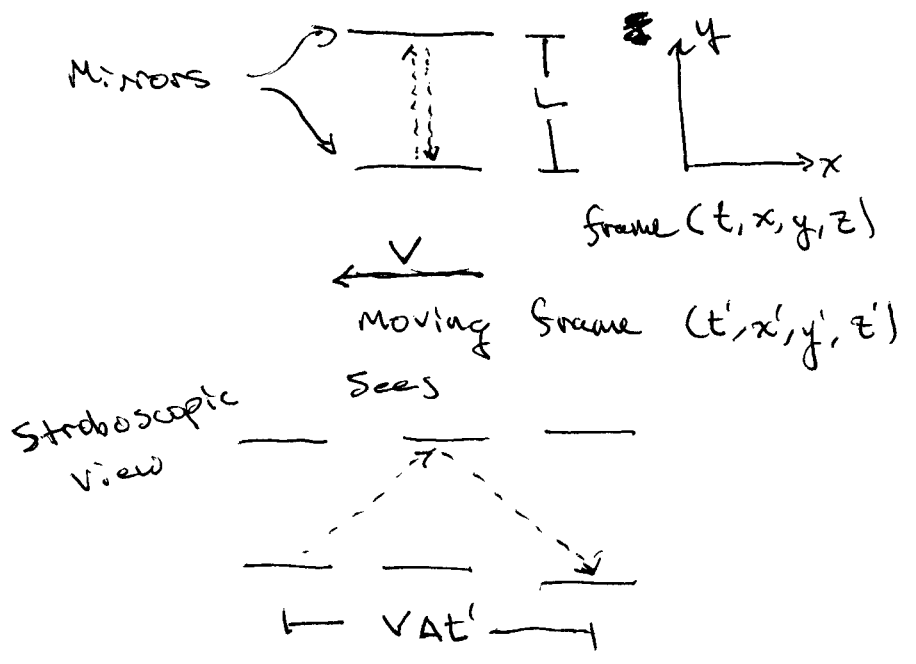
What is its geometry?
i.e. ds^2 ?

Turn our earlier observation about coordinates around:

Seek a distance all observers agree upon; an invariant distance.

Invariant under constant speed (S.R.) changes of frame.

Motivating Example: A light clock



$$= -4L^2 = -(c\Delta t)^2 + \Delta x^2 + \Delta y^2 + \Delta z^2$$

Thus

$$(\Delta S)^2 = -(c\Delta t)^2 + \Delta x^2 + \Delta y^2 + \Delta z^2$$

= same primed

Provides an invariant generalization

$$\text{or } dS^2 = dx^2 + dy^2 + dz^2$$

$$dS^2 = -c^2 dt^2 + dx^2 + dy^2 + dz^2$$

Flat spacetime line element

In (t, x, y, z) ,

$$\Delta t = \frac{2L}{c} \quad \Delta x = \Delta y = \Delta z = 0$$

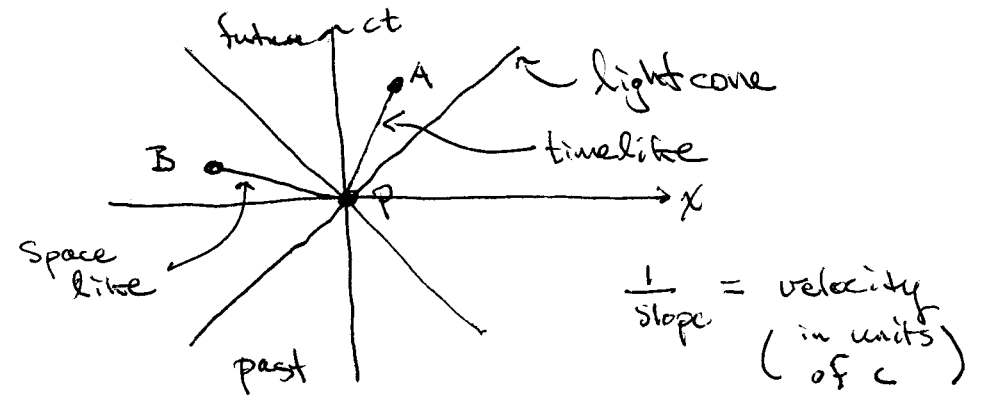
while in (t', x', y', z') check it!

$$\Delta t' = \frac{2}{c} \sqrt{L^2 + \left(\frac{\Delta x'}{2}\right)^2}, \quad \Delta x' = v\Delta t'$$

$$\Delta y' = \Delta z' = 0$$

These two frames agree on

$$-(c\Delta t')^2 + \Delta x'^2 + \Delta y'^2 + \Delta z'^2 = -4 \left[L^2 + \frac{\Delta x'^2}{4} \right] + \Delta x'^2$$

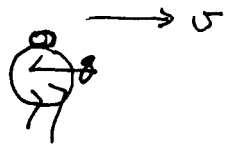


$(\Delta S)^2 > 0$ spacelike separated

$(\Delta S)^2 = 0$ null or lightlike "

$(\Delta S)^2 < 0$ timelike "

Proper Time:



particle

"proper time"
on particle's own
watch, τ



Wall Clock
Ordinary
"coordinate-
time", t

We will show briefly that

$$d\tau^2 = - ds^2 / c^2$$

length
in moving frame $L = L_0 / \gamma$
rest length

$$\gamma \equiv \frac{1}{\sqrt{1 - v^2/c^2}}$$

Moving rods are contracted (shorter)

(ii) Time Dilation

Moving clocks run slow

$$dt = \gamma d\tau$$

$$v^{x'} = \frac{v^x - v}{1 - v v^x / c^2}$$

$$\gamma = \frac{1}{\sqrt{1 - v^2/c^2}}$$

$$v^{y'} = \frac{v_y / \gamma}{1 - v v^x / c^2}$$

$$v^{z'} = \frac{v_z / \gamma}{1 - v v^x / c^2}$$

(iii) Velocity Addition

$$V_{AE} = \frac{V_{AB} + V_{BC}}{1 + \frac{V_{AB} V_{BC}}{c^2}} \quad \left(\begin{array}{l} \text{along} \\ \text{boost} \end{array} \right)$$

More specifically,

Thus for timelike curves = world lines ^{P3/5}
proper time is a nice, Lorentz
invariant, measure of "distance"
along the curve.

II Recap Relativistic Effects

(i) Lorentz Contraction

Moving rods are contracted (shorter)

All three effects (i), (ii), and (iii)
follow from the relativity of simultaneity!
(See N.D. Mermin "It's about time")

Lorentz Transformations:

Here and in what follows we adopt units in which $c=1$, see Hartle's excellent Appendix A.

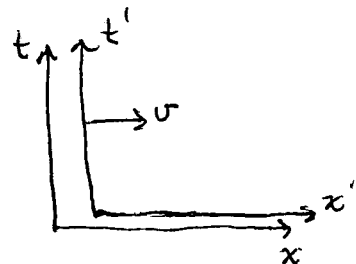
The basic connection b/w frames is

$$t' = \gamma(t - v x)$$

$$x' = \gamma(x - vt)$$

$$y' = y$$

$$z' = z$$



Origins coincide at $t=t'=0$

and view these as the components of a vector \underline{x} (I'll adopt Hartle's suggestion and use an under twiddle for 4-vectors).

4-vectors are vectors:

$$\alpha(\underline{a} + \underline{b}) = \alpha \underline{a} + \alpha \underline{b}$$

For calculating we need to express

4-vectors in a basis (recall

III Four Vectors

P4/5

So far our notation doesn't reflect our recognition of the unification of space and time.

Lorentz transformation suggests we collect

$$x^M = (x^0, x^1, x^2, x^3)$$

$$\equiv (t, x, y, z)$$

$$\underline{x} = x \hat{x} + y \hat{y} + z \hat{z}$$

$$\underline{x} = a^M \underline{e}_M$$

$$= a^t \underline{e}_t + a^x \underline{e}_x + a^y \underline{e}_y + a^z \underline{e}_z$$

Here we adopt ^{the} Einstein summation convention: repeated indices — one up, one down — are summed.

We also write

$$\underline{a} = (a^t, a^i), \quad \underline{a} = (a^t, \vec{a})$$

Scalar Product: Scalar product of basis vectors determines all scalar products, because,

$$\begin{aligned} \underline{a} \cdot \underline{b} &= (a^\mu e_{\underline{\mu}}) \cdot (b^\nu e_{\underline{\nu}}) \\ &= a^\mu b^\nu e_{\underline{\mu}} \cdot e_{\underline{\nu}} \\ &\equiv a^\mu b^\nu \eta_{\mu\nu} \end{aligned}$$

$$ds^2 = \eta_{\mu\nu} dx^\mu dx^\nu$$

and thought of this way, we call $\eta_{\mu\nu}$ the metric of flat spacetime (or Minkowski metric).

Example:

$$\begin{aligned} ds^2 &= \eta_{\mu\nu} dx^\mu dx^\nu \\ &= -dt^2 + dx^2 + dy^2 + dz^2 \\ &= -dt^2 + v^2 dt^2 \end{aligned}$$

$$\eta_{\mu\nu} \equiv e_{\underline{\mu}} \cdot e_{\underline{\nu}}$$

To find $\eta_{\mu\nu}$ we require

$$(\Delta s)^2 = \Delta \underline{x} \cdot \Delta \underline{x}$$

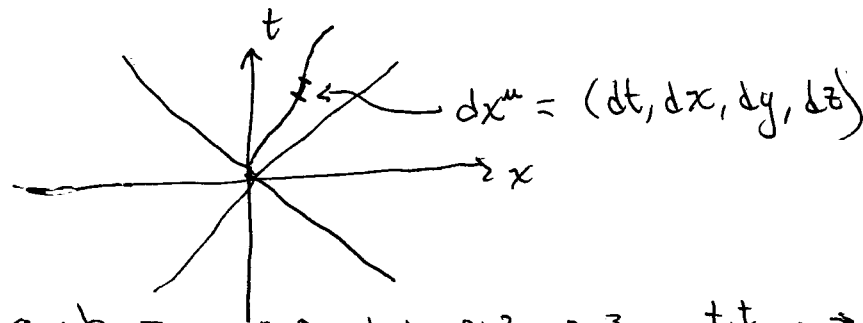
but then

$$\eta_{\mu\nu} = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

Clearly then

$$\begin{aligned} &= -(1 - v^2) dt^2 \\ &= -\frac{dt^2}{\gamma^2} = -d\tau^2 \end{aligned}$$

as mentioned before
 $d\tau^2 = -ds^2$ (in old units $= -ds^2/c^2$)



In general: $\underline{a} \cdot \underline{b} = -a^0 b^0 + a^1 b^1 + a^2 b^2 + a^3 b^3 = -\underline{a}^t \cdot \underline{b}^t + \underline{a} \cdot \underline{b}$