

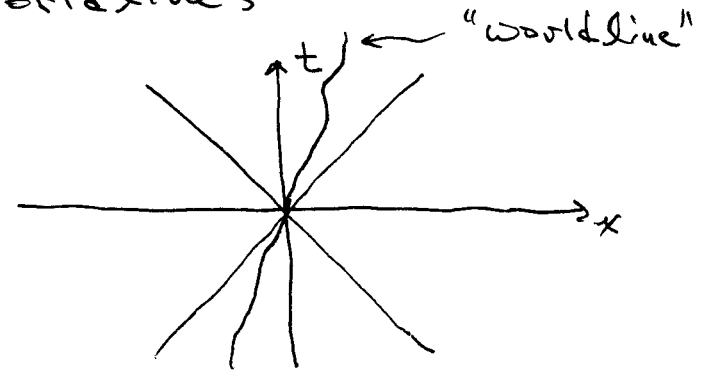
Today's Outline

I S.R. Kinematics & Dynamics

Lecture 4
Jan 26th, 2012

I S.R. Kinematics & Dynamics PL/6

Real particles follow
timelike trajectories called
worldlines



II Light

Kinematics:
Parametrize a worldline

$$x^\alpha = x^\alpha(\tau) = (t(\tau), x(\tau), y(\tau), z(\tau))$$

Eq. by proper time. Then,

Proper Velocity:

$$u^\alpha = \frac{dx^\alpha}{d\tau}$$

Why choose proper time?

$$u^t = \frac{dt}{d\tau} = \gamma = \frac{1}{\sqrt{1-\vec{v}^2}}$$

$$u^x = \frac{dx}{d\tau} = \frac{dx}{dt} \frac{dt}{d\tau} = \gamma v^x$$

etc.

So,

$$u^\alpha = (\gamma, \gamma \vec{v})$$

Important calculation,

$$\begin{aligned} \underline{u} \cdot \underline{u} &= \eta_{\alpha\beta} u^\alpha u^\beta = -\gamma^2 + \gamma^2 \vec{v}^2 \\ &= -\frac{(1-\vec{v}^2)}{(1-\vec{v}^2)} = -1. \end{aligned}$$

Proper Acceleration: (Newton:

$$a = \frac{d^2 \vec{x}}{dt^2} = \frac{d\vec{v}}{dt}$$

$$a^\alpha = \frac{du^\alpha}{d\tau}$$

Another important calculation:

$$\begin{aligned} \frac{d}{d\tau} (\underline{u} \cdot \underline{u}) &= \frac{d}{d\tau} (\eta_{\alpha\beta} u^\alpha u^\beta) \\ &= \eta_{\alpha\beta} a^\alpha u^\beta + \eta_{\alpha\beta} u^\alpha a^\beta \\ &= \underline{a} \cdot \underline{u} + \underline{u} \cdot \underline{a} = 2 \underline{a} \cdot \underline{u} \end{aligned}$$

$$\underline{p}^\alpha = m u^\alpha$$

Energy-momentum four-vector

$$\underline{p}^\alpha = (E, \vec{p})$$

Relativistic Energy:

$$E = m u^0 = m \gamma = \frac{m}{\sqrt{1-\vec{v}^2}}$$

$$\begin{aligned} &\approx m \left(1 + \frac{1}{2} \vec{v}^2 + \dots \right) \\ &= m + \frac{1}{2} m \vec{v}^2 + \dots \end{aligned}$$

Note Also:
 $\frac{\vec{p}}{E} = \frac{\delta m \vec{v}}{\delta m} = \vec{v}$
useful!

but also,

P2/6

$$\frac{d}{d\tau} (\underline{u} \cdot \underline{u}) = \frac{d}{d\tau} (-1) = 0$$

$$\Rightarrow \underline{a} \cdot \underline{u} = 0!$$

(four)-acceleration is perp.
to (four-) velocity!

Energy & Momentum:

Natural guess is correct:

Relativistic Momentum: $\vec{p} = \gamma m \vec{v} = \frac{m \vec{v}}{\sqrt{1-\vec{v}^2}}$

Important calculation:

$$\underline{p}^2 = \underline{p} \cdot \underline{p} = m^2 \underline{u} \cdot \underline{u} = -m^2$$

But then,

$$\underline{p} \cdot \underline{p} = \eta_{\alpha\beta} p^\alpha p^\beta = -E^2 + \vec{p}^2$$

$$\Rightarrow E = (m^2 + p^2)^{1/2}$$

Dynamic S:

Newton's First law carries over in the form,

$$\frac{d\vec{u}}{d\tau} = 0$$

\Rightarrow straight line motion when there are no forces and $\vec{v} = \text{const.}$

Newton's 2nd law:

$$\vec{f} = m \frac{d\vec{u}}{d\tau} = m \vec{a} = \frac{d\vec{p}}{d\tau}$$

it is useful to ^{also} introduce the three-force

$$\vec{F} = \frac{d\vec{p}}{dt} \leftarrow \begin{array}{l} \text{relativistic} \\ \text{momentum} \\ \text{coordinate time} \end{array}$$

In terms of three-force we have

$$\vec{f} = \frac{d\vec{p}}{d\tau} = \frac{d\vec{p}}{dt} \frac{dt}{d\tau} = \gamma \vec{F}$$

Also, $\vec{f} \cdot \vec{u} = 0 \Rightarrow -f^t \gamma + \gamma^2 \vec{F} \cdot \vec{v} = 0$
 $\Rightarrow f^t = \gamma \vec{F} \cdot \vec{v}$

and we call \vec{f} the P3/6
four-force. (or Minkowski force).

Note that,

$$\vec{f} \cdot \vec{u} = m \vec{a} \cdot \vec{u} = 0$$

So only three of the equations $\vec{f} = m \vec{a}$ are independent.

Despite its hybrid character

so,
$$\vec{f} = (\gamma \vec{F} \cdot \vec{v}, \gamma \vec{F})$$

Note that

$$\frac{dP^0}{d\tau} = \frac{dP^0}{dt} \frac{dt}{d\tau} = \gamma \frac{dE}{dt} = f^0 = \gamma \vec{F} \cdot \vec{v}$$

$$\Rightarrow \frac{dE}{dt} = \vec{F} \cdot \vec{v}$$

Relativistic version of power — which we know follows from other E.O.M.

Example: Find $x(t)$ for motion in 1 dimension under a constant "three"-force F .



Say it starts at $x=0$, from rest at $t=0$. Well,

$$\vec{F} = \frac{d\vec{p}}{dt}$$

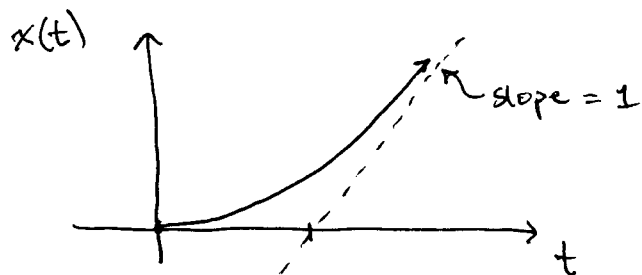
$$\Rightarrow dp = F dt \Rightarrow p = Ft$$

$$\Rightarrow \gamma m v = Ft \Rightarrow \frac{m v}{\sqrt{1-v^2}} = Ft$$

Separate and integrate again,

$$\int_0^x dx' = \int_0^t \frac{F t'}{\sqrt{m^2 + F^2 t'^2}} dt'$$

$$\Rightarrow x(t) = \frac{1}{F} \sqrt{m^2 + F^2 t^2} - \frac{m}{F}$$



Solve for V ,

P4/6

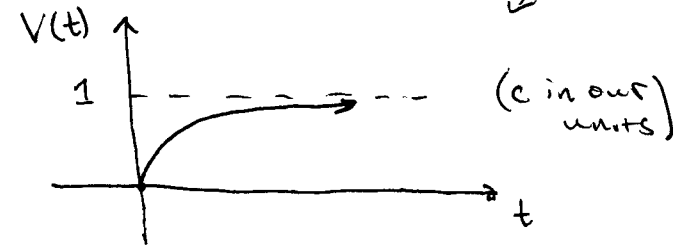
$$\frac{m^2 V^2}{1-V^2} = F^2 t^2$$

$$\Rightarrow m^2 V^2 = F^2 t^2 - F^2 t^2 V^2$$

$$\Rightarrow V^2 (m^2 + F^2 t^2) = F^2 t^2$$

$$\Rightarrow V = \frac{Ft}{\sqrt{m^2 + F^2 t^2}}$$

Plot



II Light

Light has the amazing property

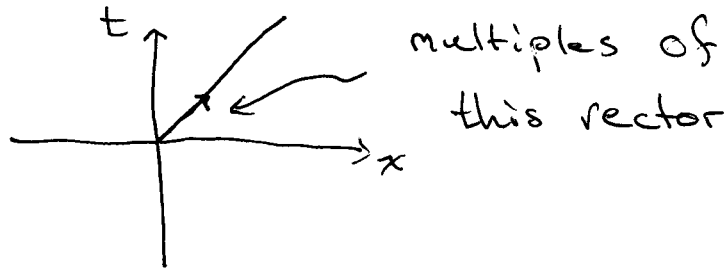
$$m=0$$

and travels at $V=1$. We've seen that

$$ds^2 = -d\tau^2 = 0$$

for null separations. This means we can't use s or τ to parametrize light trajectories.

Instead invent a parametrization - e.g. for $x=t$ we take



just like a free particle.

~~if~~ instead $x^\alpha = \tilde{u}^\alpha \sigma^3$, then same trajectory,

$$u^\alpha = \frac{dx^\alpha}{d\sigma} = (3\sigma^2, 3\sigma^2, 0, 0)$$

and $\underline{u} \cdot \underline{u} = 0$ still but

$$\frac{d^2 x^\alpha}{d\sigma^2} = \frac{du^\alpha}{d\sigma} = 6\sigma \tilde{u}^\alpha \neq 0.$$

We call a parametrization of a

i.e.

PS/6

$$x^\alpha = \tilde{u}^\alpha \lambda, \text{ with } \tilde{u}^\alpha = (1, 1, 0, 0)$$

Note that

$$u^\alpha = \frac{dx^\alpha}{d\lambda} = \tilde{u}^\alpha$$

and

$$\underline{\tilde{u}} \cdot \underline{\tilde{u}} = -1 + 1 + 0 + 0 = 0,$$

i.e. $\underline{\tilde{u}}$ is a null vector.

Also,

$$\frac{d^2 x^\alpha}{d\lambda^2} = \frac{du^\alpha}{d\lambda} = 0$$

light ray affine if it gives the free particle eg. for light rays. Affine parameters are preferred, but not unique.

Energy & Momentum

Einstein recognized that for a photon,

$$E = h\nu$$

Planck's const. / 2π \nearrow \nwarrow freq.

The Energy-momentum relation

$$E^2 = m^2 + \vec{p}^2$$

continues to hold for light w/ $m=0$,

$$\Rightarrow |\vec{p}| = E$$

But then,

$$\vec{p} = \hbar \vec{k} \quad \text{w/} \quad |\vec{k}| = \omega$$

and \hat{k} in direction of motion. So,

$$\text{light: } p^\alpha = (E, \vec{p}) = \hbar(\omega, \vec{k}) \equiv \hbar k^\alpha$$

so,

$$\omega = \gamma(\omega' - v k^{x'})$$

But $k^{x'} = \omega' \cos \alpha'$, so

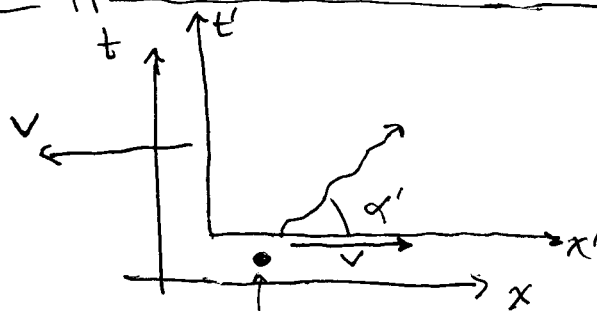
$$\omega = \gamma(\omega' - v \omega' \cos \alpha')$$

$$\Rightarrow \omega' = \omega \frac{\sqrt{1-v^2}}{1-v \cos \alpha'}$$

 Doppler Effect

k^α is the wave 4-vector. P6/6

Doppler Shift and Relativistic



light source moving in (t', x')

k^α is a 4-vector, transforms under Lorentz trans. just like x^α