

Outline

I Last time

II Energy & Momentum
in Relativity

III Conservation of

relativistic energy/mom.

IV Massless particles

$$\boxed{\begin{aligned} (x^0)' &= \gamma (x^0 - \beta x^1) \\ (x^1)' &= \gamma (x^1 - \beta x^0) \\ (x^2)' &= x^2 \\ (x^3)' &= x^3 \end{aligned}}$$

Proper time: particle's wrist

watch time, call it τ .

$$dt = \gamma d\tau$$

Proper velocity

$$v = \frac{dx}{d\tau} = \gamma \frac{dx}{dt} = \gamma v$$

or more generally

$$v^\mu = \frac{dx^\mu}{d\tau} = \gamma (c, v_x, v_y, v_z).$$

Modern

Day 5

I

Lorentz transformations

$$(x^0)' = \gamma (x^0 - \beta x^1)$$

$$(x^1)' = \gamma (x^1 - \beta x^0)$$

$$(x^2)' = x^2$$

$$(x^3)' = x^3$$

Note that

$$\gamma = \sqrt{1 - \frac{1}{\gamma^2} \frac{v^2}{c^2}} \Rightarrow \gamma^2 = \frac{1}{1 - \frac{1}{\gamma^2} \frac{v^2}{c^2}}$$

$$\Rightarrow \gamma^2 \left(1 - \frac{1}{\gamma^2} \frac{v^2}{c^2}\right) = 1 \Rightarrow \gamma^2 = 1 + \frac{v^2}{c^2}$$

$$\Rightarrow \gamma = \sqrt{1 + \frac{v^2}{c^2}}$$

II Momentum: classically: mass-velocity

So

$$\vec{p} = m \vec{v} \quad \text{or} \quad \vec{p} = m \vec{\gamma} ?$$

The 1st def. would \Rightarrow conservation of momentum is not a possible law - if true in S, not true in S'.

$\vec{p}^o = m\vec{\eta}^o = m\vec{v}$

Define: Relativistic momentum:

$$\tilde{p} = m\tilde{\eta} = \gamma m\vec{v} = \frac{m\vec{v}}{\sqrt{1 - v^2/c^2}}$$

This means that \tilde{p} is the spatial part of a 4-vector:

$$p^\mu = m\eta^\mu$$

Note: E is not zero for a particle at rest

$$E = mc^2 \equiv R \quad \text{"Rest energy"}$$

$$\text{Then } T = E - R = \gamma mc^2 - mc^2 = (\gamma - 1)mc^2$$

"Kinetic ener."

$$= \left(\left(1 - \frac{v^2}{c^2} \right)^{-\frac{1}{2}} - 1 \right) mc^2 \approx \frac{1}{2} mv^2$$

(non-relativistic)

$$E = \sqrt{mc^2}$$

or

$$E = \sqrt{1 - \frac{v^2}{c^2}}$$

so,

$$p^\mu = \left(\frac{E}{c}, p_x, p_y, p_z \right)$$

"Energy-Momentum 4-vector"

Useful fact:

$$E^2 - p^2 c^2 = \frac{m^2 c^4}{1 - \frac{v^2}{c^2}} - \frac{m^2 v^2 c^2}{1 - \frac{v^2}{c^2}}$$

$$= \frac{m^2 c^4}{1 - \frac{v^2}{c^2}} \left(1 - \frac{v^2}{c^2} \right)$$

$$= m^2 c^4$$

$$\text{So, } \boxed{E^2 - p^2 c^2 = m^2 c^4}$$

gets you directly from E^2 to \tilde{p}^2 .

Question: What's the zero p_2/p_3 ("temporal") component?

III Conservation laws:

In any collision process (including decays) energy & momentum are conserved.

Example:



(before)

(after)

Question: What's the mass of the composite lump?



rest energy and mass decrease, while K.E. (kinetic energy) increases.

def. in elastic collision is one in which K.E. is conserved (in practice, some particles exit as in).

$$\text{IV } m=0$$

In classical mechanics: No such thing!
 $P = m\vec{v} = 0$, $T = \frac{1}{2}mv^2 = 0$

forget gluons, approximate for neutrinos)

$$\frac{2mc^2}{\sqrt{1 - \frac{q}{c^2}}} = M_{\text{jet}} c^2$$

$$\Rightarrow M = \frac{2}{\sqrt{1 - \frac{q}{c^2}}} m = \frac{10}{3} m = 2.5 m$$

Mass is not conserved, in general!

Note: all forms of 'internal energy' are included in the rest mass of a composite object ($m \approx M$)

Relativity: Consider! If $v=c$,

then the denominators are also zero.

$$E = \gamma c? \quad p = \gamma v c?$$

$$\text{But, } E^2 - p^2 c^2 = m^2 c^4 = 0$$

$$\Rightarrow E^2 = p^2 c^2$$

$$\Rightarrow E = \gamma c$$

In fact, they exist (photons,