

## Outline

I Where are we?

Where are we going?

II Central Forces: Reduction

Mechanics

Day 13

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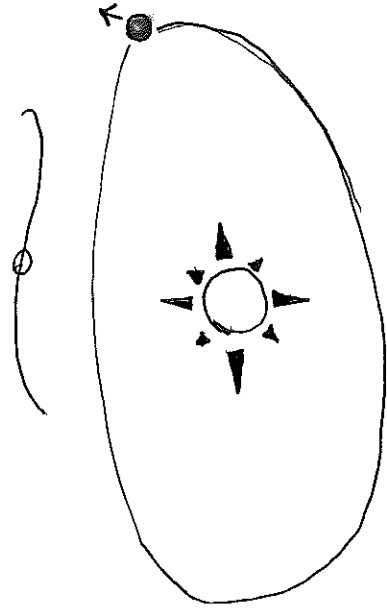
I We've now touched on all three of the great advantages of Lagrangian mechanics:

- Can use generalized coord.s
- Completely eliminate the forces of constraint.
- Marvelous dual role of symmetry.

In all of physics, but also for illustrating the power of reduction.

symmetry  $\Leftrightarrow$  cons. law

The second facet of this advantage is the ability to remove D.O.F. altogether. We call this process reduction. We turn now to one of the most important examples



$\delta S = 0$

is just another way of saying that  $S$  is stationary for the physical path (similar to  $df=0$  in calculus). But we know what the conditions for

$S = \int_{t_1}^{t_2} \mathcal{L}(q_i, \dot{q}_i, t) dt$   
to be stationary are:

$\frac{\partial \mathcal{L}}{\partial q_i} = \frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \dot{q}_i} \right) \quad (i=1,2)!$

The constraints have been completely

III Central Forces: Reduction

We've setup the general formalism, now, let's apply it! Major application: the two body problem. — two bodies that experience an interaction force, between one another, but no external forces.

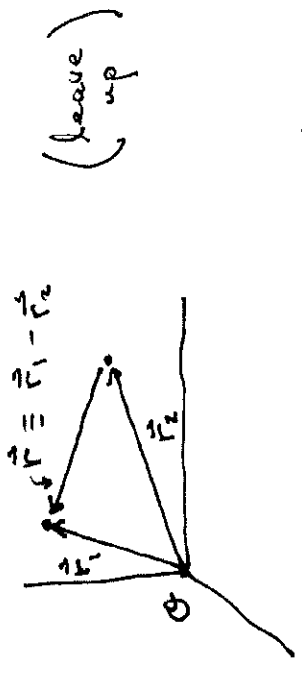
Examples:  $U = - \frac{G m_1 m_2}{|\vec{r}_1 - \vec{r}_2|}$  gravitation

or  $U = \frac{k q_1 q_2}{|\vec{r}_1 - \vec{r}_2|}, \quad k = \frac{1}{4\pi\epsilon_0}$  Coulomb force



We've done it. Lagrangian mechanics provides a completely independent formulation of mechanics that is equivalent to Newton's. From here forward we can use whatever formulation is most convenient. I will do this!

3D space



Our present (remarkable) goal is Reduction: show that these problems of 6 D.O.F. are equivalent to a 1 D.O.F. problem. This is not due to constraints; instead we will use symmetry and conservation laws (Noether's ideas).

Useful Definition: A central force is one for which the force is always directed toward or away from a fixed "force center." If we take the center as the origin,  $\theta$ , then these are ~~fixed and for external~~ radial forces: magnitude; can be + or - and depends on  $\vec{r}$

$$F(\vec{r}) = f(r) \hat{r}$$



If we assume our central force is derivable from a potential,  $\vec{F} = -\nabla U$ , we have an interesting

We'll often drop 'conservative' but this isn't a great practice. End of useful definition.

In both examples above  $U$  only depends on  $|\vec{r}_1 - \vec{r}_2| \equiv |\vec{r}| = r$  and we will call  $\vec{F} \equiv \vec{F}_1 - \vec{F}_2$  the "relative position". This dependence is a manifestation of the translational invariance of these problems; Mathematically,

$$\vec{r}_1 \rightarrow \vec{r}_1 + \vec{E} \quad \text{and} \quad \vec{r}_2 \rightarrow \vec{r}_2 + \vec{E}$$

gives rise to exactly the same potential and hence force btwn the two particles.

consequence for  $f(\vec{r})$ . In spherical coordinates

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$$\vec{F} = -\nabla U = -\frac{\partial U}{\partial r} \hat{r} - \frac{1}{r} \frac{\partial U}{\partial \theta} \hat{\theta} - \frac{1}{r \sin \theta} \frac{\partial U}{\partial \phi} \hat{\phi}$$

but  $\vec{F}$  is central, so,  $U$  is spher. symmetric:  
 $\frac{\partial U}{\partial \theta} = 0$  and  $\frac{\partial U}{\partial \phi} = 0 \Rightarrow U = U(r) = U(|\vec{r}|)$   
 therefore

$$f(\vec{r}) \equiv -\frac{\partial U(r)}{\partial r} = f(|\vec{r}|) = f(r)$$

and  $\vec{F} = f(r) \hat{r}$  (conservative central force)

Noether's thm.  $\Rightarrow$  there is a corresponding conserved quantity. Here we can't translate the particles independently but only both together and this corresponds to conservation of total momentum. (check it using Noether's thm!)

This all indicates that we want conglomerate general coordinates in addition to our relative ones  $\vec{r}$ .

Introduce the CM (center of mass) position:

$$\vec{R} = \frac{m_1 \vec{r}_1 + m_2 \vec{r}_2}{m_1 + m_2} = \frac{m_1 \vec{r}_1 + m_2 \vec{r}_2}{M}$$

Then

$$\begin{aligned} \vec{r}_i &= \vec{R} + \frac{m_j}{M} \vec{r} \\ &= \frac{(m_1 + m_2) \vec{r}_i}{M} = \vec{r}_i \quad \checkmark \end{aligned}$$

check it

and

$$\vec{r}_2 = \vec{R} - \frac{m_1}{M} \vec{r}$$

mass and a convenient shorthand is

$$\mu = \frac{m_1 m_2}{m_1 + m_2}$$

We call  $\mu$  the "reduced mass." Then

$$L_{rel} = \frac{1}{2} \mu \dot{r}^2 - U(r)$$

is a Lagrangian that only depends on  $\vec{r}$  and  $\dot{\vec{r}}$ .

Let's look at E.O.M.s. For  $\vec{R}$  we have

$$\frac{\partial L}{\partial \vec{R}_i} = \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{\vec{R}}_i} \right) = M \ddot{\vec{R}}_i \quad (i=1,2,3)$$

Put this all into the (conservative) central force Lagrangian,

$$\begin{aligned} L &= \frac{1}{2} m_1 \dot{\vec{r}}_1^2 + \frac{1}{2} m_2 \dot{\vec{r}}_2^2 - U(r) \\ &= \frac{1}{2} M \dot{\vec{R}}^2 + \left( \frac{1}{2} \frac{m_1 m_2}{M} \dot{\vec{r}}^2 - U(r) \right) \\ &\equiv L_{cm} + L_{rel} \end{aligned}$$

The quantity  $\frac{m_1 m_2}{M}$  has units of

or

$$M \ddot{\vec{R}} = 0.$$

This implies  $M \dot{\vec{R}} = m_1 \dot{\vec{r}}_1 + m_2 \dot{\vec{r}}_2$

$= \vec{p}_1 + \vec{p}_2 = \vec{p}_{tot}$  is constant!

The center of mass moves as if it were a free particle (at a constant velocity).

Idea: Change reference frame to one in which  $\dot{\vec{R}} = 0$  and  $\vec{R} = 0$  (CM at origin) then

$$L = L_{rel}$$

Now,  $L$  only depends on  $\vec{r}$  so we're down to 3D.O.F.

There is still a rotational sym.:  $|\vec{R}(\vec{r}_1 - \vec{r}_2)| = |\vec{r}_1 - \vec{r}_2|$   
 Then, Noether's theorem implies that  
 the total angular momentum is conserved!

$$\vec{L} = \vec{r}_1 \times \vec{p}_1 + \vec{r}_2 \times \vec{p}_2$$

In the CM frame  $\vec{R} = 0$  and  $\dot{\vec{R}} = 0$ , so,

$$\vec{r}_1 = \frac{m_2}{M} \vec{r} \quad \text{and} \quad \vec{r}_2 = -\frac{m_1}{M} \vec{r}$$

and

$$\begin{aligned} \dot{\vec{L}} &= \frac{m_2}{M} \dot{\vec{r}} \times m_1 m_2 \dot{\vec{r}} + \left(-\frac{m_1}{M}\right) \dot{\vec{r}} \times m_2 \left(-\frac{m_1}{M}\right) \dot{\vec{r}} \\ &= \frac{m_1 m_2}{M^2} (m_2 \dot{\vec{r}} \times \dot{\vec{r}} + m_1 \dot{\vec{r}} \times \dot{\vec{r}}) = \dot{\vec{r}} \times \mu \dot{\vec{r}} \end{aligned}$$

By convention we choose the Z-axis to be parallel to  $\vec{L}$ . Then all of the motion lies in the xy-plane (i.e.  $\vec{r}$  and  $\dot{\vec{r}}$  lie in plane perp. to  $\vec{L}$ ).

We have reduced the problem to one in 2 D.O.F. at this point.

Now choose polar coordinates in the xy-plane  $(r, \phi)$ . Our Lagrangian is,

$$\mathcal{L} = \frac{1}{2} \mu (\dot{r}^2 + r^2 \dot{\phi}^2) - U(r).$$

The vector  $\vec{L}$  is constant, P56 which means both its magnitude and its direction are constant.

Because  $\dot{\vec{L}} = \dot{\vec{r}} \times \mu \dot{\vec{r}}$  it must be perpendicular to both  $\vec{r}$  and  $\dot{\vec{r}}$ . As long as  $\dot{\vec{L}} \neq 0$ , the two vectors  $\vec{r}$  and  $\dot{\vec{r}}$  span a plane.

Convention: We are free to orient our axes w/in the CM frame however we like.

We notice that  $\phi$  is ignorable

$$\Rightarrow \frac{\partial \mathcal{L}}{\partial \phi} = \text{const.}$$

$$\Rightarrow \mu r^2 \dot{\phi} = \text{const} \equiv L \quad (\text{Eq. 1})$$

This is just the z-component of angular momentum that we knew we had around (recall  $\mu r^2 = I$  for pt. particle and  $\dot{\phi} = \omega = \text{angular speed}$ ,  $L_z = I\omega$ ). Initial conditions set the value of  $L$  and so we

can use (Eq. 1) to get rid of  $\phi$ ,

$$(Eq. 1) \Rightarrow \dot{\phi} = \frac{L}{\mu r^2}$$

once we're working with the E.O.M.

This will complete our reduction,

all that remains is 1 D.O.F.,  $r$ .

Reduction has been a major success for central forces; we've achieved the

best possible outcome: 6 D.O.F.  $\rightarrow$  1 D.O.F.