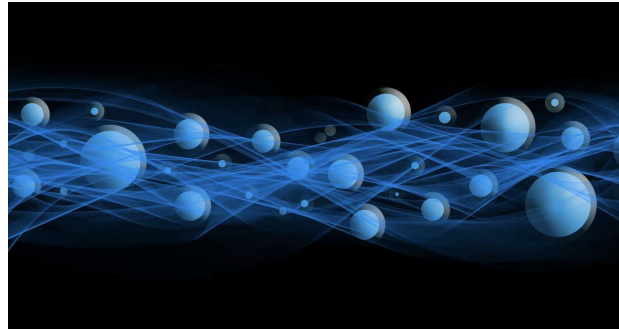




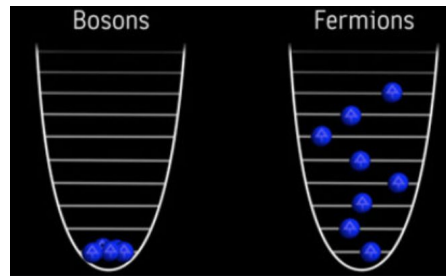
Bose-Einstein Condensation



Boson:

a particle (such as a photon or meson) whose spin quantum number **is** zero or an integral number

These can be packed into the same state



Photons, Phonons:

- The Chemical potential μ is zero
- The number of Photons in a given system change frequently

$$N_{total} \neq \text{constant}$$

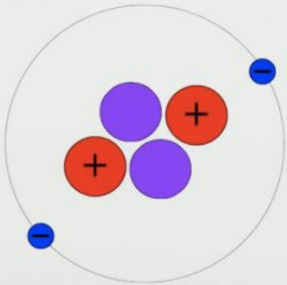
$$\mu = 0$$



“Ordinary Bosons”

For example Helium-4 has bosonic while Helium-3 has Fermionic statistic

Helium-4

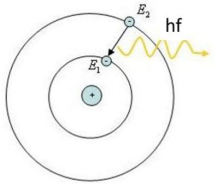


These Bosons are atoms whose fundamental particles add up to an integer spin

The total number of particles is fixed

$$N_{total} = \text{Constant}$$

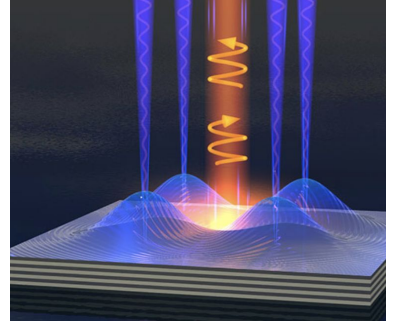
Bose-Einstein Condensation occurs when a gas of Bosons will abruptly condense into the ground state as the Temperature goes below a certain value

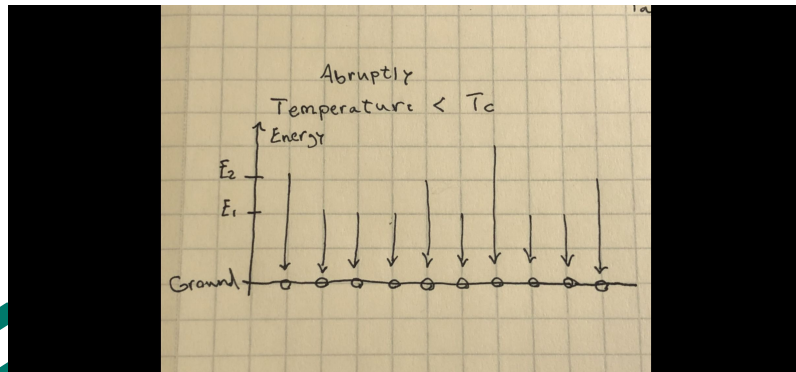
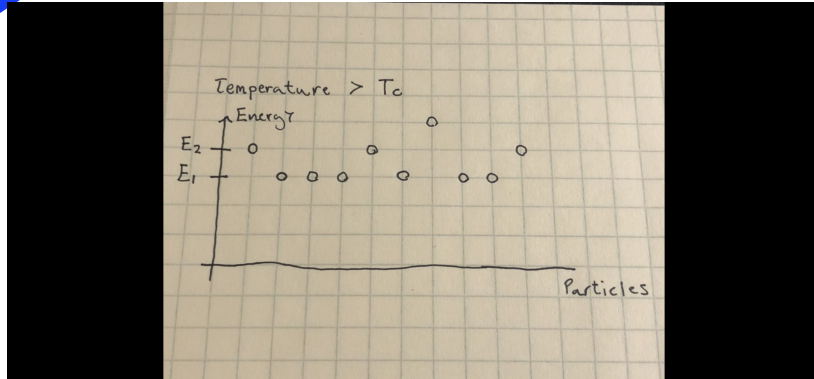


The amount of atoms in a given state is given by the **Bose-Einstein Distribution:**

For an “Ordinary Boson” the chemical potential (rather than being fixed at zero) is now a nontrivial function of the density and temperature.

$$N_0 = \frac{1}{e^{\frac{(\epsilon_0 - \mu)}{kT}} - 1}$$





It is simplest to consider the
limit $T \rightarrow 0$

When the temperature is zero Kelvin, all
the atoms will be in the lowest-energy
available state.

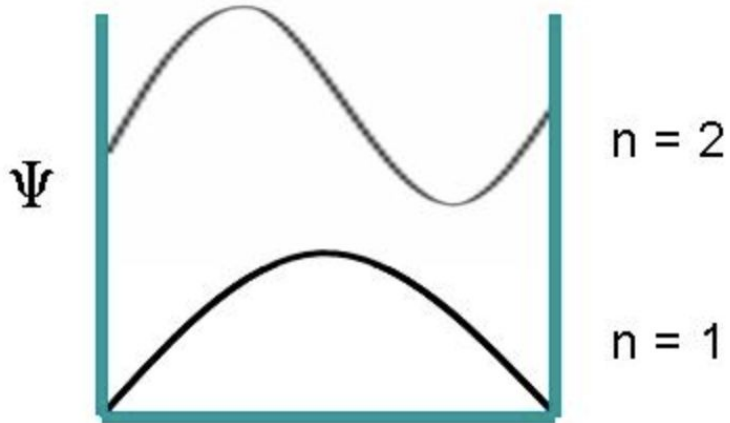
Since arbitrarily many bosons are allowed
in any state, this means that every atom
will be in the ground state.

(Here when I simply say “State” I mean a single particle state.)

The average number of atoms in a state is given by the Bose-Einstein distribution
So the average number of atoms in the ground state is given by the B.E. distribution when epsilon is equal to the energy of the ground state.

$$N_0 = \frac{1}{e^{\frac{(\epsilon_0 - \mu)}{kT}} - 1}$$

$\varepsilon_0 =$ *The energy of the ground state*



To find epsilon we use the same logic that we have used previously of the DeBroglie wavelength that fits in a box.

For a box of side lengths L , and volume V^3 the energy of the ground state is:

$$\varepsilon_0 = \frac{h^2}{8mL^2} (1^2 + 1^2 + 1^2) = \frac{3h^2}{8mL^2} \quad (7.118)$$

Is there a better way to understand the way that N depends on ϵ , μ and kT ?

$$N_0 = \frac{1}{e^{\frac{(\epsilon_0 - \mu)}{kT}} - 1} \quad (7.119)$$

A Taylor Series Expansion to simplify the B.E. distribution

$$N_0 = \frac{1}{e^{\frac{(\epsilon_0 - \mu)}{kT}} - 1}$$

(When $N_0 \gg 1$)

For the Taylor series to work the exponent needs to be close to zero

This happens when T is sufficiently low, N will be quite large. In this case the denominator of this expression must be small, which implies the exponential is close to 1.

This means that the exponent $\frac{\epsilon_0 - \mu}{kT}$ must be small, and so our series should work

The average number of atoms in the ground state is given by
7.120

$$N_0 = \frac{1}{1 + \frac{(\epsilon_0 - \mu)}{kT} - 1} = \frac{kT}{\epsilon_0 - \mu} \quad \left(\text{When } N_0 \gg 1 \right) \quad (7.120)$$

What does this mean?

$$N_0 = \frac{kT}{\epsilon_0 - \mu}$$

When $T=0$, we know that the chemical potential μ is equal to ϵ .

When the temperature is non-zero, but still very small so that most of the atoms are in the ground state, μ is a tiny bit less than ϵ .

When $T=0$, $\mu = \epsilon_0$

and when $T > 0$, but still $\ll 1$ $\mu < \epsilon_0$

How low must the temperature be, in order for N_0 to be large

The general condition that determines μ is that the sum of the Bose-Einstein distribution over all states must add up to the total number of atoms, N :

$$N = \sum_{\text{all } s} \frac{1}{e^{\frac{\epsilon_s - \mu}{kT}} - 1} \quad (7.121)$$

To solve this we could keep guessing values of μ until the sum works perfectly and repeat the process for each value of T .

Converting to an integral

$$N = \int_0^{\infty} g(\epsilon) \frac{1}{e^{\frac{(\epsilon - \mu)}{kT}} - 1} \quad (7.122)$$

Valid when $kT \gg \epsilon_0$

So that the number of terms that contribute significantly to the sum is large.

This integral contains a new function $g(\epsilon)$.

$$N_0 = \frac{1}{e^{\frac{(\varepsilon_0 - \mu)}{kT}} - 1}$$



General Condition for μ

$$N = \sum_{\text{all } s} \frac{1}{e^{\frac{\epsilon_0 - \mu}{kT}} - 1}$$

$$\text{---->} \quad N = \int_0^{\infty} g(\epsilon) \frac{1}{e^{\frac{(\epsilon - \mu)}{kT}} - 1}$$

The function $g(\epsilon)$ is the **density of states**: The number of single-particle states per unit energy.

- For spin-zero Bosons
- Confined in a box of Volume V

Bear with me here

Same as electrons

$$g(\epsilon) = \frac{2}{\sqrt{\pi}} \left(\frac{2\pi m}{h^2} \right)^{\frac{3}{2}} V \sqrt{\epsilon}$$

$$N = \int_0^{\infty} g(\epsilon) \frac{1}{e^{\frac{(\epsilon - \mu)}{kT}} - 1}$$

Plugging in for μ

$$N = \frac{2}{\sqrt{\pi}} \left(\frac{2\pi m}{h^2} \right)^{\frac{3}{2}} V \int_0^{\infty} \frac{\sqrt{\epsilon} d\epsilon}{e^{\frac{\epsilon}{kT}} - 1}$$

With a substitution of variables

$$x = \frac{\epsilon}{kT}$$

$$N = \frac{2}{\sqrt{\pi}} \left(\frac{2\pi m k T}{h^2} \right)^{\frac{3}{2}} V \int_0^{\infty} \frac{\sqrt{x} dx}{e^x - 1}$$

This integral over x gives 2.315;
Combining this number with the
factor

$$\frac{2}{\sqrt{\pi}}$$

We get the formula

$$N = 2.612 \left(\frac{2\pi m k T}{h^2} \right)^{\frac{3}{2}} V \quad (7.125)$$

This formula can only work for a
specific Temperature that we will
call T_c .

$$N = 2.612 \left(\frac{2\pi m k T_c}{h^2} \right)^{\frac{3}{2}} V \quad (7.126)$$

Or

$$k T_c = 0.527 \left(\frac{h^2}{2\pi m} \right) \left(\frac{N}{V} \right)^{\frac{2}{3}}$$

Now to prove that we made a good choice of μ

$$N = 2.612 \left(\frac{2\pi m k T}{h^2} \right)^{\frac{3}{2}} V \quad \text{Eq. (7.125)}$$

At $T > T_c$ the chemical potential must be significantly less than 0.

$$N = \int_0^{\infty} g(\epsilon) \frac{1}{e^{\frac{(\epsilon - \mu)}{kT}} - 1} \quad \text{Eq. (7.122)}$$

A negative value of μ will yield a result for N that is smaller than the right hand side of equation 7.125.

At temperatures lower than T_c it gets a bit more tricky

The bottom line is this:

At temperatures higher than T_c the chemical potential is negative and essentially all of the atoms are in the excited state. At temperatures lower than T_c , the chemical potential is very close to zero and the number of atoms in excited states is given by equation 7.127

$$N_{excited} = \left(\frac{T}{T_c} \right)^{\frac{3}{2}} N \text{ when } T < T_c$$