## **Today**

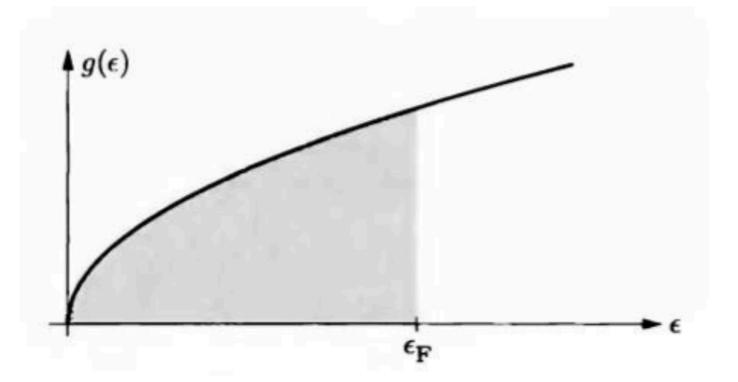
- I. Announcements: One final class on Monday of completion days (Dec. 14th), Final due Friday Dec. 18th at 10pm ET
- II. Last Time
- III. Density of States Wrap Up
- IV. Balthazar's Guest Lecture on Debye Theory of Solids
- V. CAFE forms for Thermal: <a href="https://tools.bard.edu/tools/cafeform/">https://tools.bard.edu/tools/</a> <a href="mailto:cafeform/">cafeform/</a>
- I. Bruno told us about Bose-Einstein condensation. At low temperatures all Bosons drop into the lowest energy state and they behave collectively in quantum mechanical way.

Introduced the idea of the density of states: divide energy space into bands, then the density of states tells me how many quantum states there are in a given band of energies.

Let's recast our computation of the energy of the Fermi gas in terms of the energies of the states:

$$\epsilon = \frac{h^2}{8mL^2}n^2$$
, gives  $n = \sqrt{\frac{8mL^2}{h^2}}\sqrt{\epsilon}$ , and so  $dn = \sqrt{\frac{8mL^2}{h^2}}\frac{1}{2\sqrt{\epsilon}}d\epsilon$ 

$$U = \pi \int_{0}^{n_{max}} \frac{h^{2}n^{2}}{8mL^{2}} n^{2} dn = \int_{0}^{\epsilon_{F}} \epsilon \left[ \frac{\pi}{2} \left( \frac{8mL^{2}}{h^{2}} \right)^{3/2} \sqrt{\epsilon} \right] d\epsilon$$



Let's recast our computation of the energy of the Fermi gas in terms of the energies of the states:

$$\epsilon = \frac{h^2}{8mL^2}n^2$$
, gives  $n = \sqrt{\frac{8mL^2}{h^2}}\sqrt{\epsilon}$ , and so  $dn = \sqrt{\frac{8mL^2}{h^2}}\frac{1}{2\sqrt{\epsilon}}d\epsilon$ 

$$U = \pi \int_{0}^{n_{max}} \frac{h^{2}n^{2}}{8mL^{2}} n^{2} dn = \int_{0}^{\epsilon_{F}} \epsilon \left[ \frac{\pi}{2} \left( \frac{8mL^{2}}{h^{2}} \right)^{3/2} \sqrt{\epsilon} \right] d\epsilon$$

Note that we can write this in one more way:

$$g(\epsilon) = \frac{\pi (8m)^{3/2}}{2h^3} V \sqrt{\epsilon}.$$

By definition we can also compute the number of particles in our

system using this: 
$$N = \int_0^{\epsilon_F} g(\epsilon) d\epsilon$$
  $(T = 0)$ .

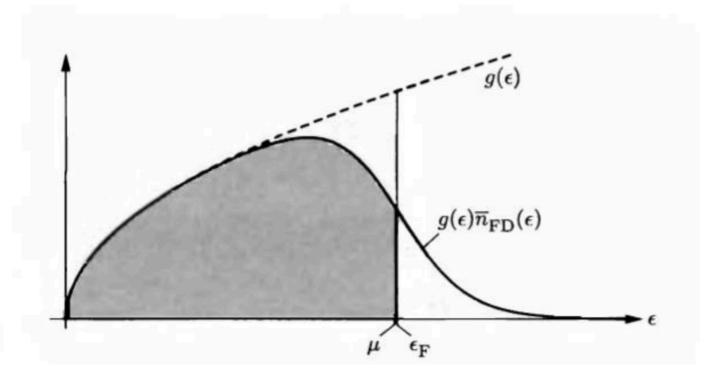
By definition we can also compute the number of particles in our

system using this:  $N = \int_{0}^{\epsilon_F} g(\epsilon) d\epsilon$  (T = 0).

We can do this at nonzero temperatures, as long as we take the probability of occupation into account

$$N = \int_0^\infty g(\epsilon) \overline{n}_{FD}(\epsilon) d\epsilon = \int_0^\infty g(\epsilon) \frac{1}{e^{(\epsilon - \mu)/kT} + 1} d\epsilon$$

$$\mu(T) \neq \epsilon_{\rm F}$$
 except when  $T = 0$ .



By definition we can also compute the number of particles in our system using this:  $N = \int_0^{\epsilon_F} g(\epsilon) d\epsilon$  (T = 0).

We can do this at nonzero temperatures, as long as we take the probability of occupation into account

$$N = \int_{0}^{\infty} g(\epsilon) \overline{n}_{FD}(\epsilon) d\epsilon = \int_{0}^{\infty} g(\epsilon) \frac{1}{e^{(\epsilon - \mu)/kT} + 1} d\epsilon$$

$$U = \int_{0}^{\infty} \epsilon g(\epsilon) \overline{n}_{FD}(\epsilon) d\epsilon = \int_{0}^{\infty} \epsilon g(\epsilon) \frac{1}{e^{(\epsilon - \mu)/kT} + 1} d\epsilon$$