

Today

I last time

II Measurement

III Light's Energy & Momentum

General Relativity

Day 11

Feb 24th, 2016 P/3

I. Newton:

1st $\frac{d\mathbf{u}}{dt} = 0 \iff$

Straight line motion
 $\vec{v} = \text{const.}$

2nd $\frac{d\mathbf{u}}{dt} = m \frac{d\mathbf{u}}{d\tau} = m \mathbf{a} = \frac{d\mathbf{p}}{d\tau}$

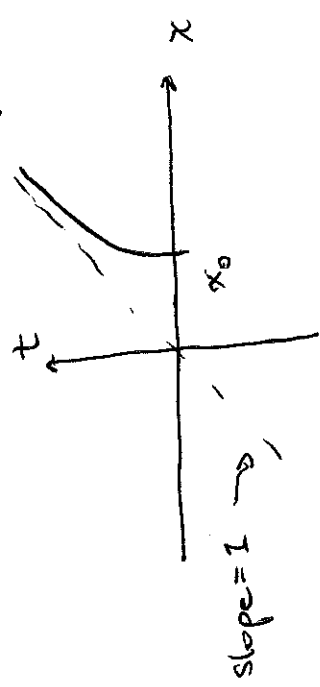
3 4-force or Minkowski force

3-force: $\vec{F} = \frac{d\vec{p}}{dt}$

$f^\mu = (\gamma \vec{F} \cdot \vec{v}, \gamma \vec{F})$

Ex: Constant 3-force in 1D:

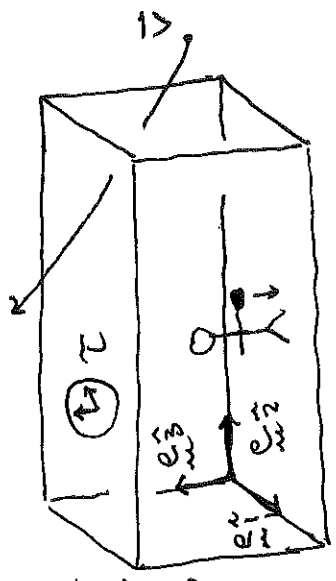
$[x(t) - x_0 + \frac{m}{F}]^2 = \frac{m^2}{F^2} + t^2$



$$V(t) = \frac{Ft}{\sqrt{m^2 + F^2 t^2}}$$

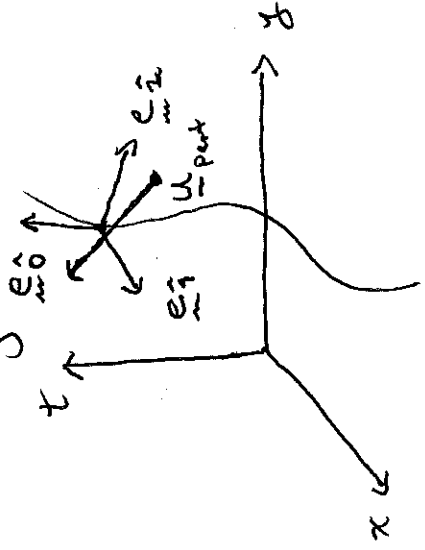
II The relational aspect of both S.R. and G.R. can make measurements difficult to interpret. However, there is a very general strategy that is also quite natural operationally.

Observers make local measurements by setting up axes and clocks:



four orthogonal unit vectors $e_{\hat{1}}, e_{\hat{2}}, e_{\hat{3}}, e_{\hat{0}}$

Spacetime diagram



Recall $u_{obs} \cdot u_{obs} = -1$, so

u_{obs} is a unit (timelike) vector.

So an observer naturally chooses

The $\hat{p}^{\tilde{\alpha}}$ represent the energy (\hat{p}^0) and momentum (\hat{p}^i) as measured by our moving observer. To calculate these components we use

$$\begin{aligned} \hat{p}^{\tilde{\alpha}} \cdot \hat{p} &= \hat{p}^{\tilde{\alpha}} \cdot \hat{e}_{\tilde{\alpha}} \cdot \hat{p} \\ &= \hat{p}^{\tilde{\alpha}} \eta_{\tilde{\alpha}\tilde{\beta}} \hat{p}^{\tilde{\beta}} \end{aligned}$$

assume these are orthonormal

$$\Rightarrow \hat{p}^{\tilde{0}} = -\hat{e}_{\tilde{0}} \cdot \hat{p}, \quad \hat{p}^{\tilde{i}} = \hat{e}_{\tilde{i}} \cdot \hat{p}, \text{ etc.}$$

$$\hat{e}_{\tilde{0}} = u_{obs}$$

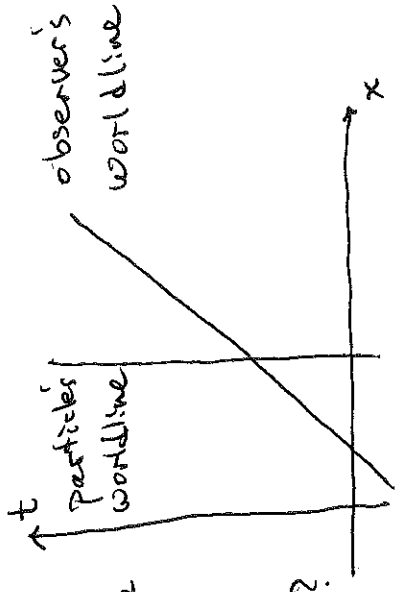
This captures the observation that we are always at rest in our own frame, but moving in time.

To determine what an observer measures we decompose, e.g., a particle's 4-momentum \hat{p} into the observer's frame:

$$\hat{p} = \hat{p}^{\tilde{\alpha}} \hat{e}_{\tilde{\alpha}}$$

Let's see it work in practice:

Ex. Consider a particle, mass m , at rest in an inertial frame. An observer moves along the x -axis of this frame with velocity \vec{v} .



What is the energy of the particle according to this observer?

(3) In the observer's frame $P^3/3$

$$u_{\text{obs}} = (1, 0, 0, 0) = \underline{e}_0$$

$$P = (\gamma m, -\gamma m v, 0, 0)$$

and so,

$$E = -\underline{e}_0 \cdot P = \gamma m.$$

Methods (2) and (3) illustrate an important point: Dot products are Lorentz invariants and so you can evaluate them in any

The energy-momentum relation

$$E^2 = m^2 + \vec{p}^2$$

Continues to hold for light w/
 $m=0, \Rightarrow \boxed{|\vec{p}| = E}$

Because $E = \hbar \omega$ it is convenient to factor out an \hbar from \vec{p} too:

$$\vec{p} = \hbar \vec{k}$$

and this implies

$$|\vec{p}| = \hbar |\vec{k}| = E = \hbar \omega$$

or

$$|\vec{k}| = \omega$$

Three different methods:
 (1) According to the observer the particle moves with velocity $-\vec{v}$

and so,

$$E = \gamma m, \quad \gamma = \gamma(|\vec{v}|).$$

(2) In the frame drawn

$$u_{\text{obs}} = (\gamma, \gamma v, 0, 0) = \underline{e}_0, \quad \gamma = \gamma(|\vec{v}|)$$

$$P = (m, 0, 0, 0)$$

and so,

$$E = -\underline{e}_0 \cdot P = -(-m\gamma) = m\gamma.$$

convenient coordinates - you will always get the same answer.
 Method (1) will become difficult in G.R. and methods (2) & (3) become far superior.

III Light has the amazing property

$m=0$ and travels at $v=c$.
 Einstein recognized that for a photon

$$E = \hbar \omega \leftarrow \text{angular frequency}$$

Planck's const. / $2\pi = 10^{34} \text{ J}\cdot\text{s}$